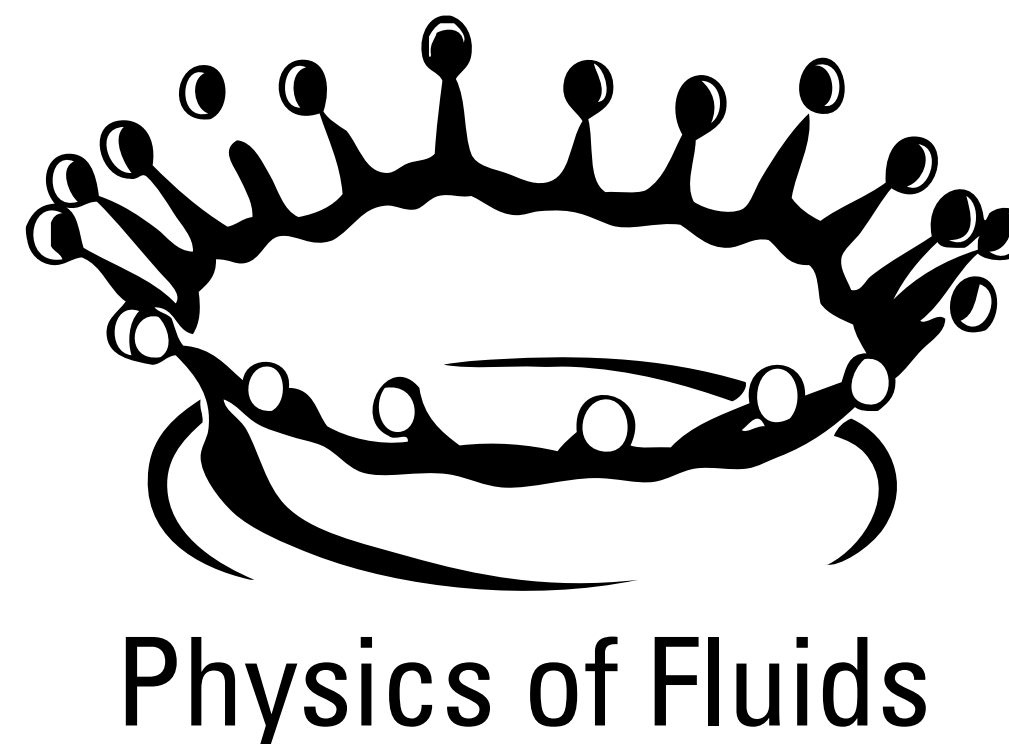
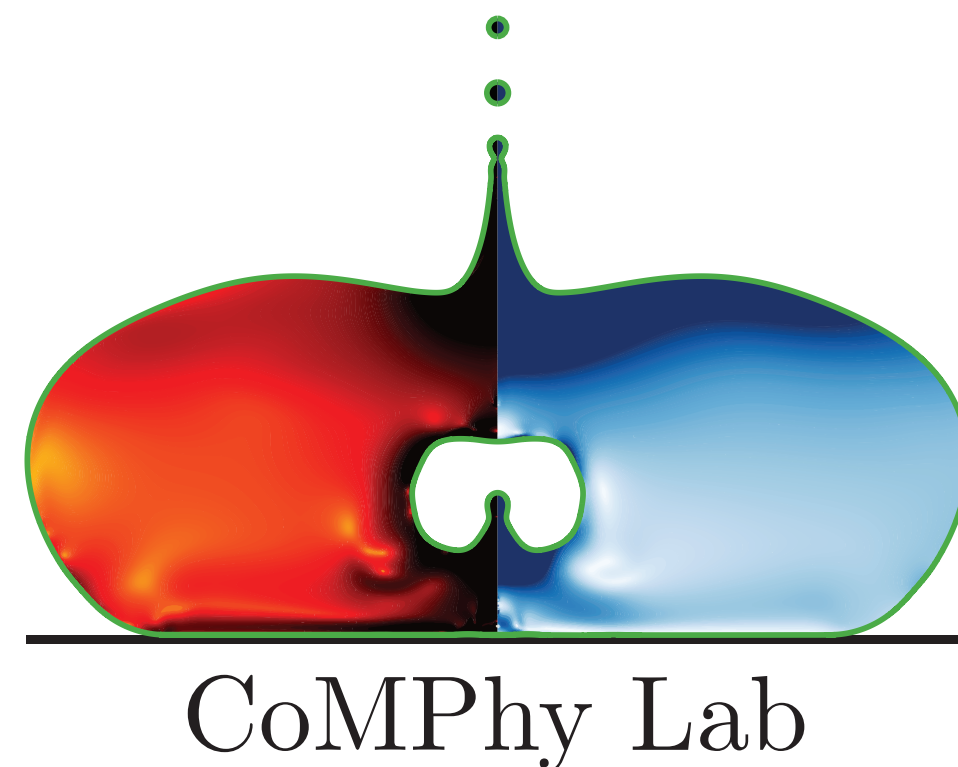


Holey Sheets

Ayush Dixit



**UNIVERSITY
OF TWENTE.**

Coauthors



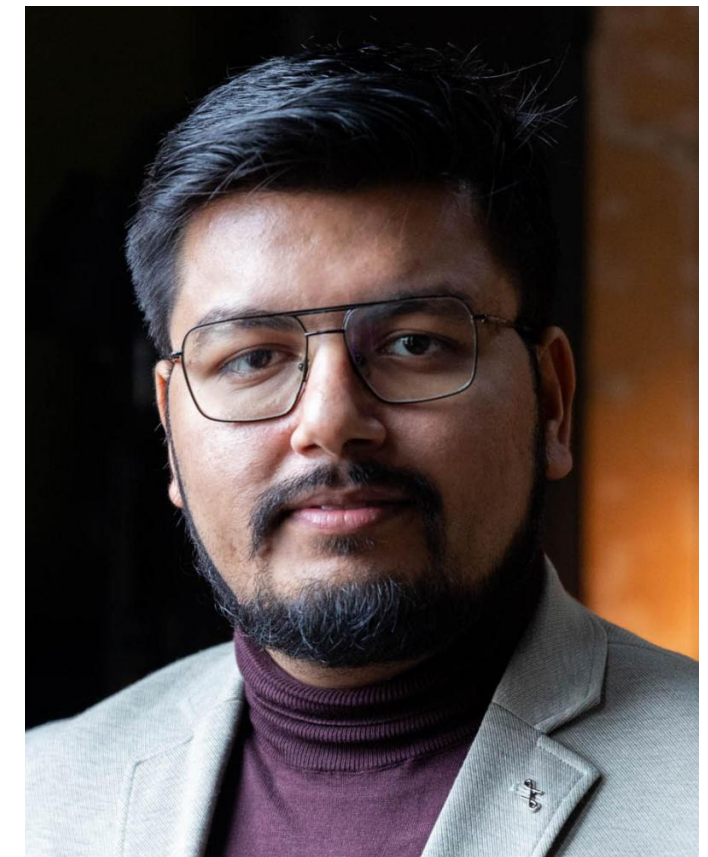
Chunheng Zhao,
CC New York



Stéphane Zaleski,
Sorbonne Université



Detlef Lohse,
Uni. Twente



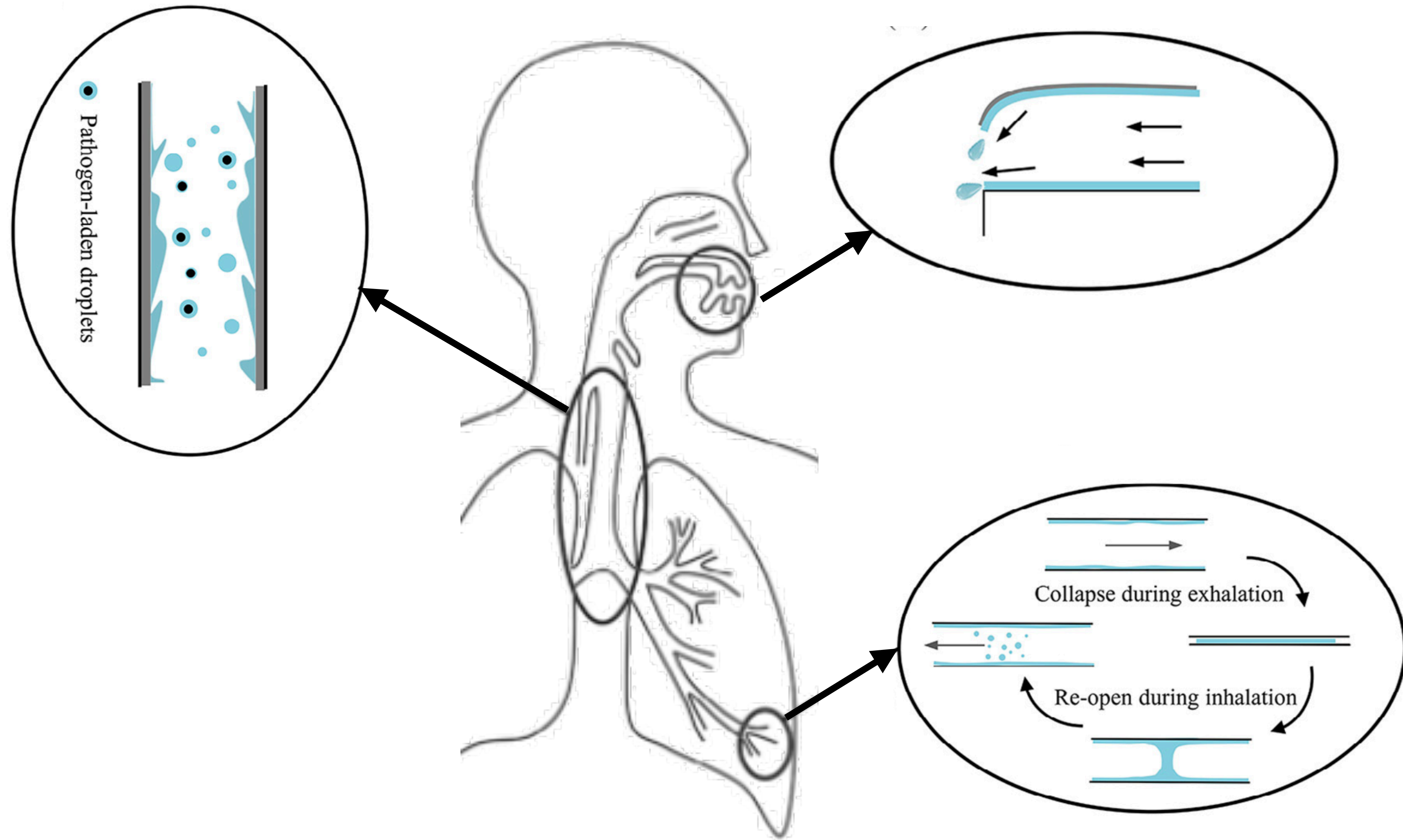
Vatsal Sanjay,
Uni. Twente,
Durham Uni.

Why do we care?



Scharfman, Techet, Bush & Bourouiba, Exp. Fluids 57, 24 (2016)

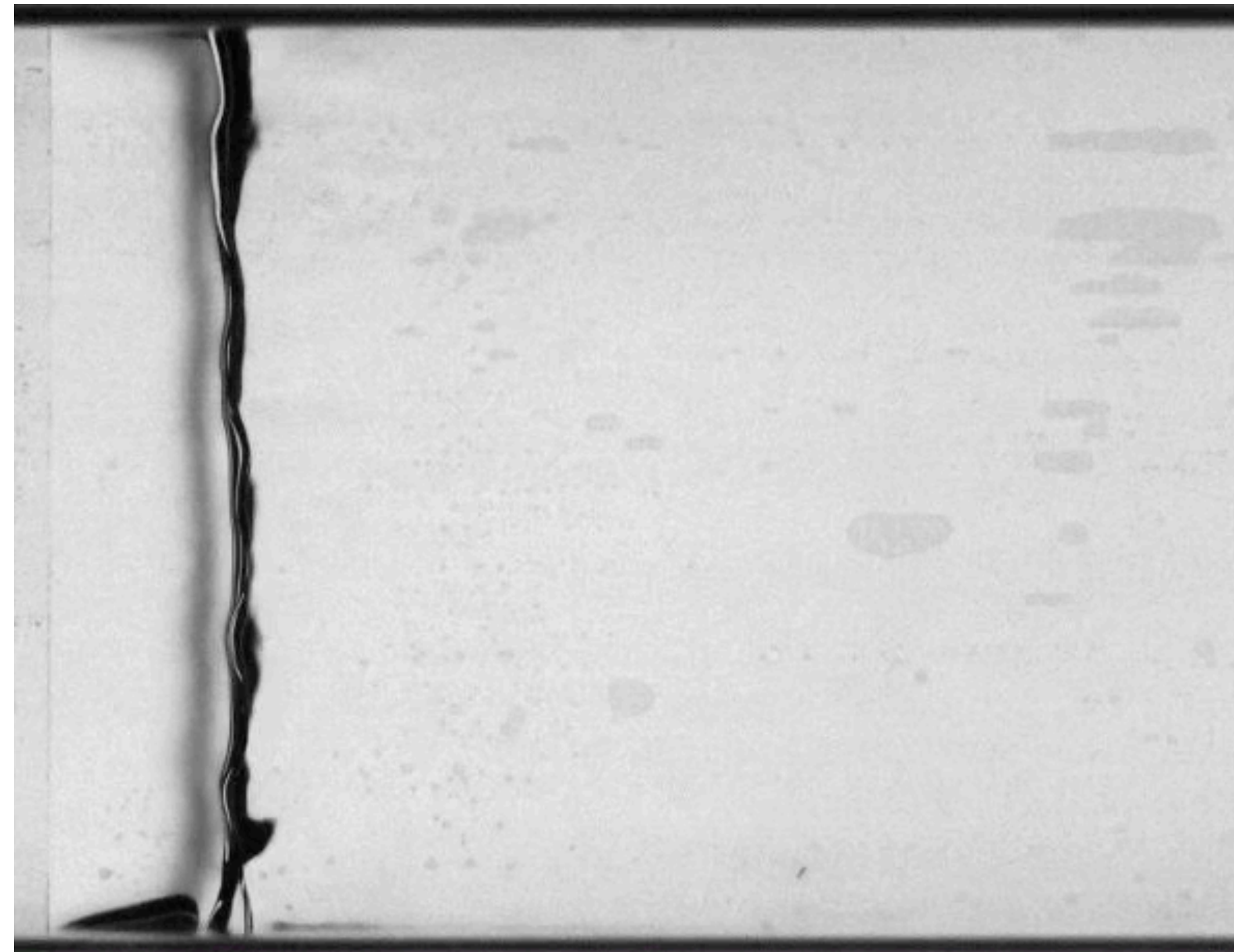
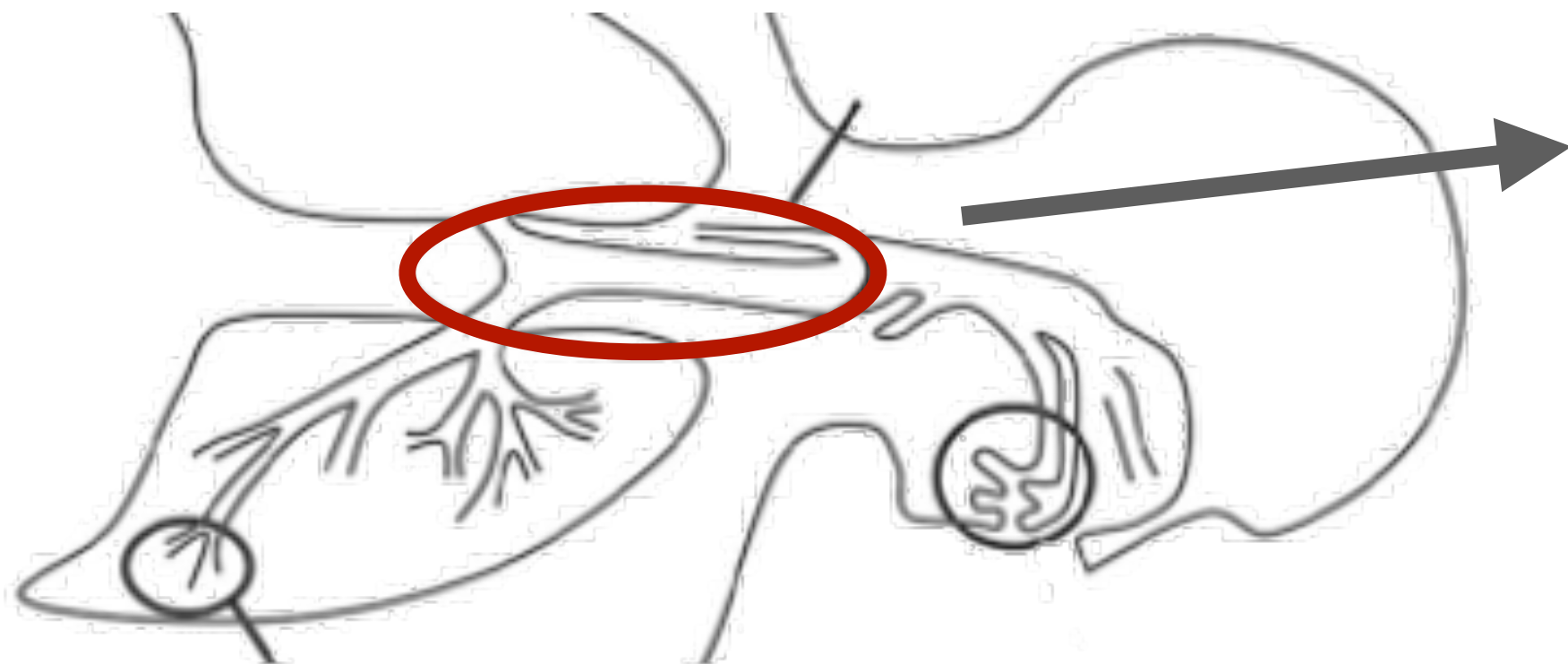
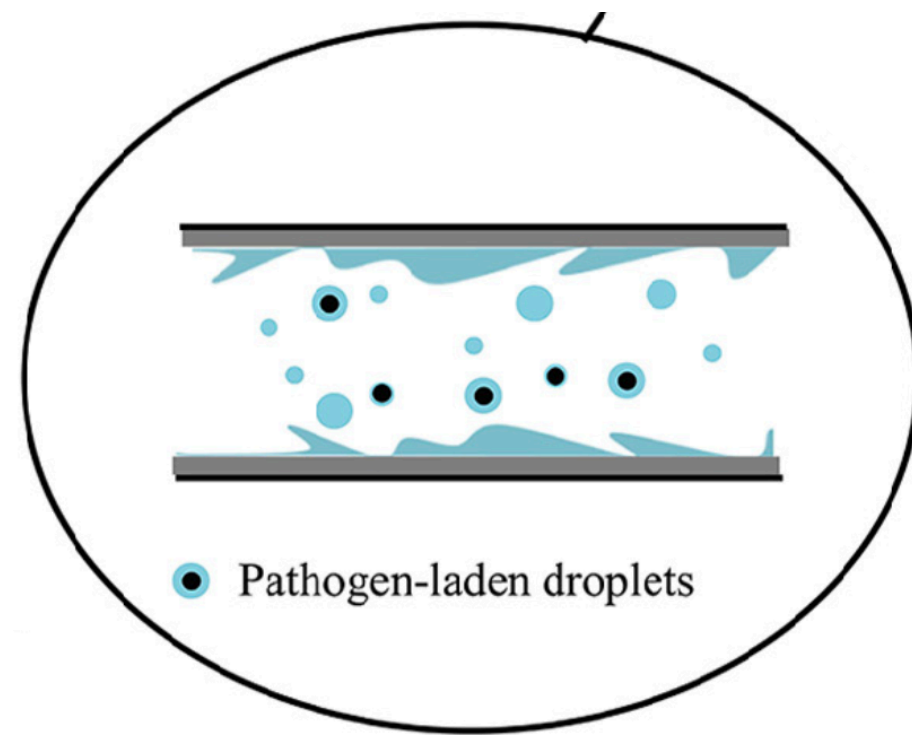
Where are the droplets coming from?



Wie & Li, Am. J. Infect. Control 44, 102 (2016)

Pöhlker, Pöhlker, Krüger et al., Rev. Mod. Phys. 95, 045001 (2023)

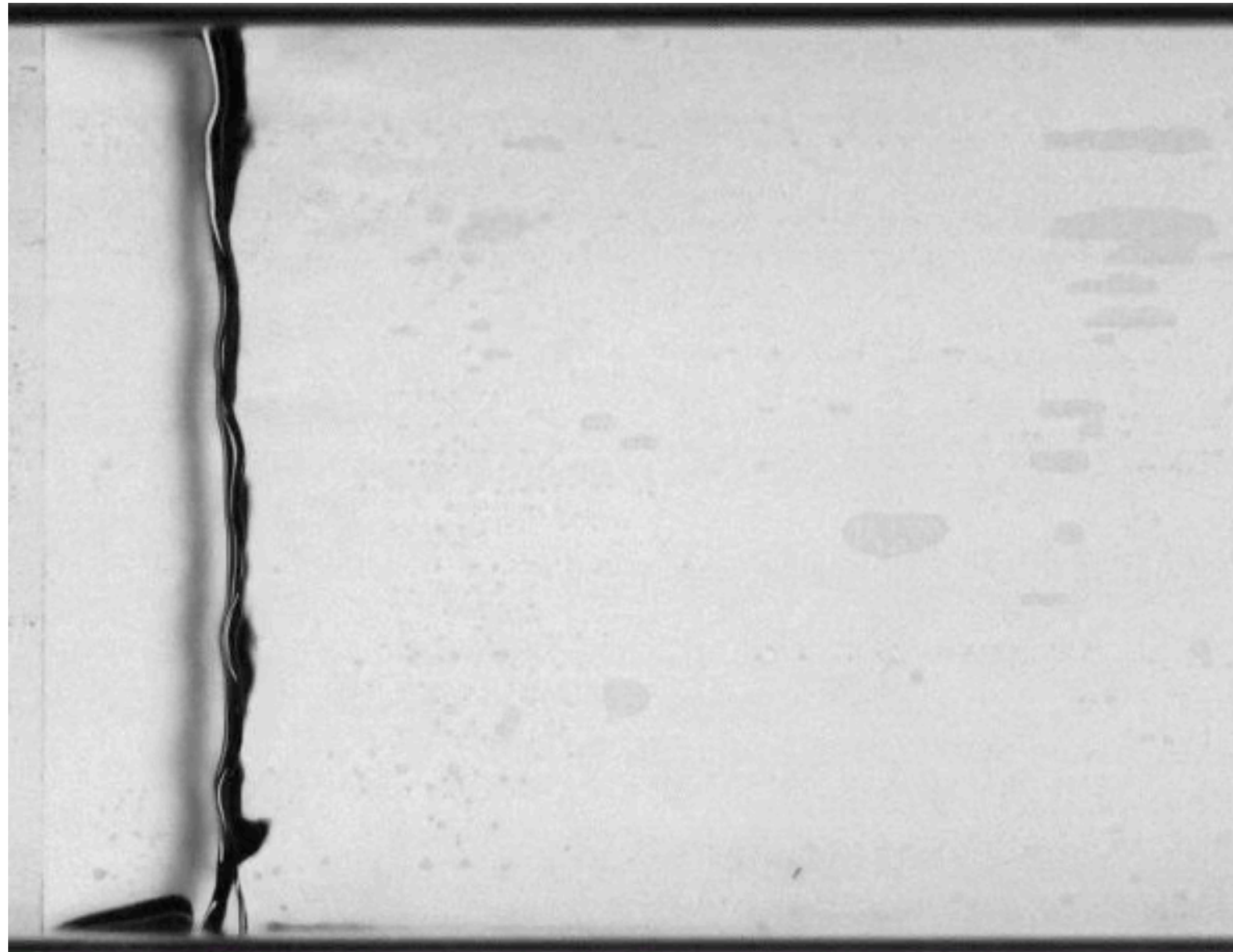
Coughing as mimicked in the lab



Kant, Pairetti, Saade, Popinet, Zaleski & Lohse, Phys. Rev. Fluids 8, 074802 (2023)

Li, Saade, Zaleski, Sen, Kant & Lohse, arXiv (2025)

Coughing as mimicked in the lab



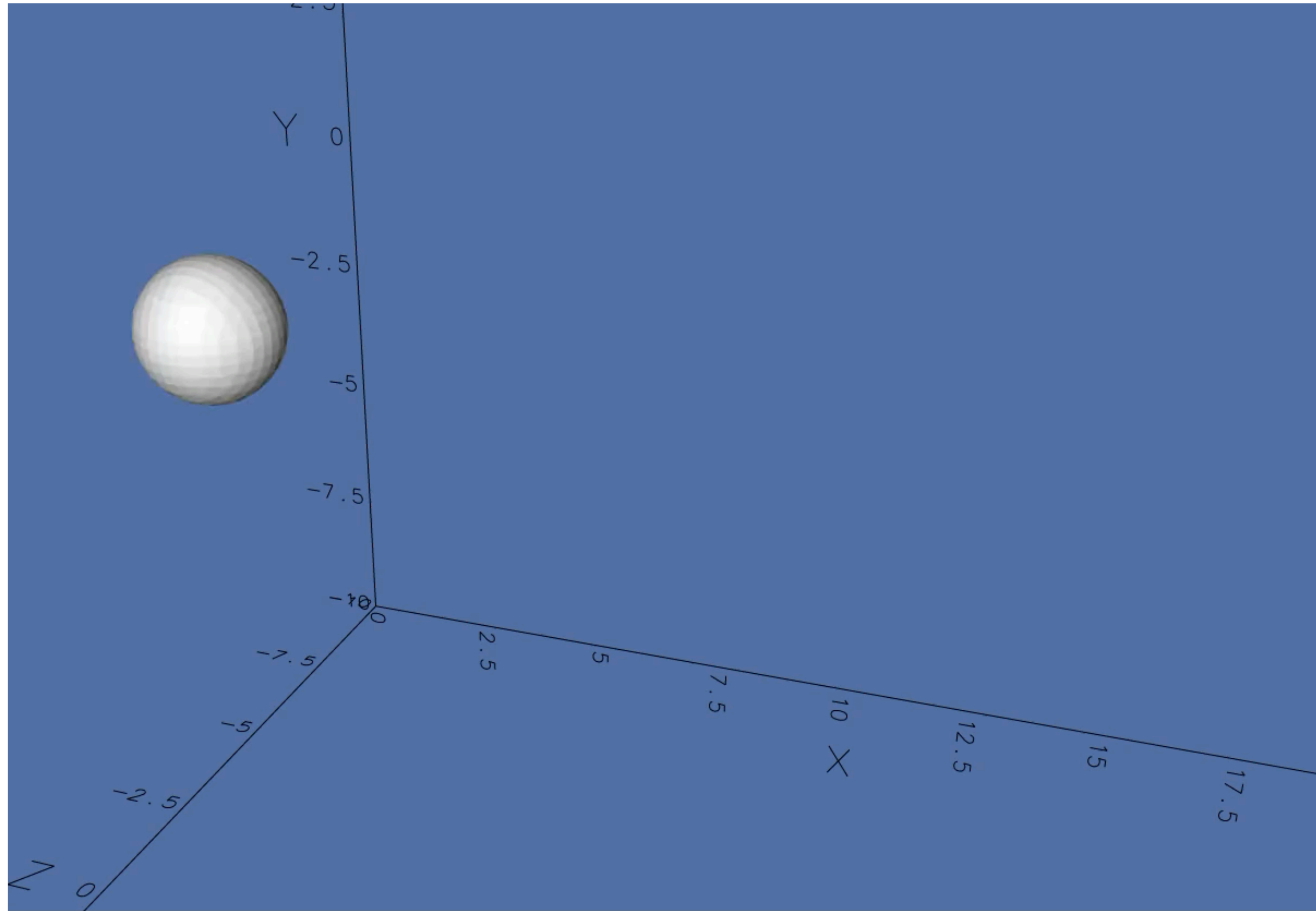
Jetting

Bag-mediated atomisation

Kant, Pairetti, Saade, Popinet, Zaleski & Lohse, Phys. Rev. Fluids 8, 074802 (2023)

Li, Saade, Zaleski, Sen, Kant & Lohse, arXiv (2025)

Bag breakup of a single droplet



Newtonian

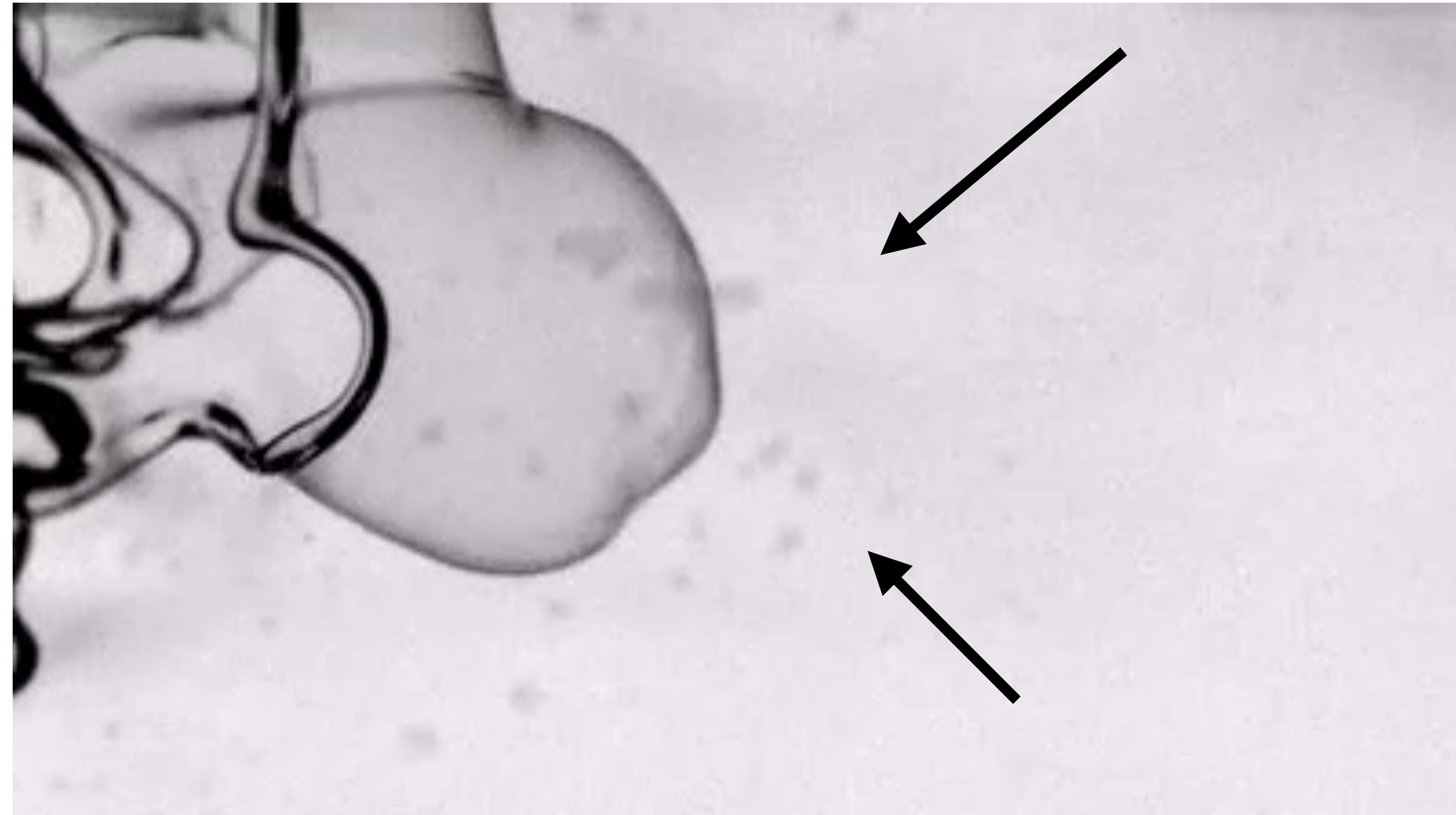
$$Oh = 3 \times 10^{-3}$$

$$We = 7.5$$

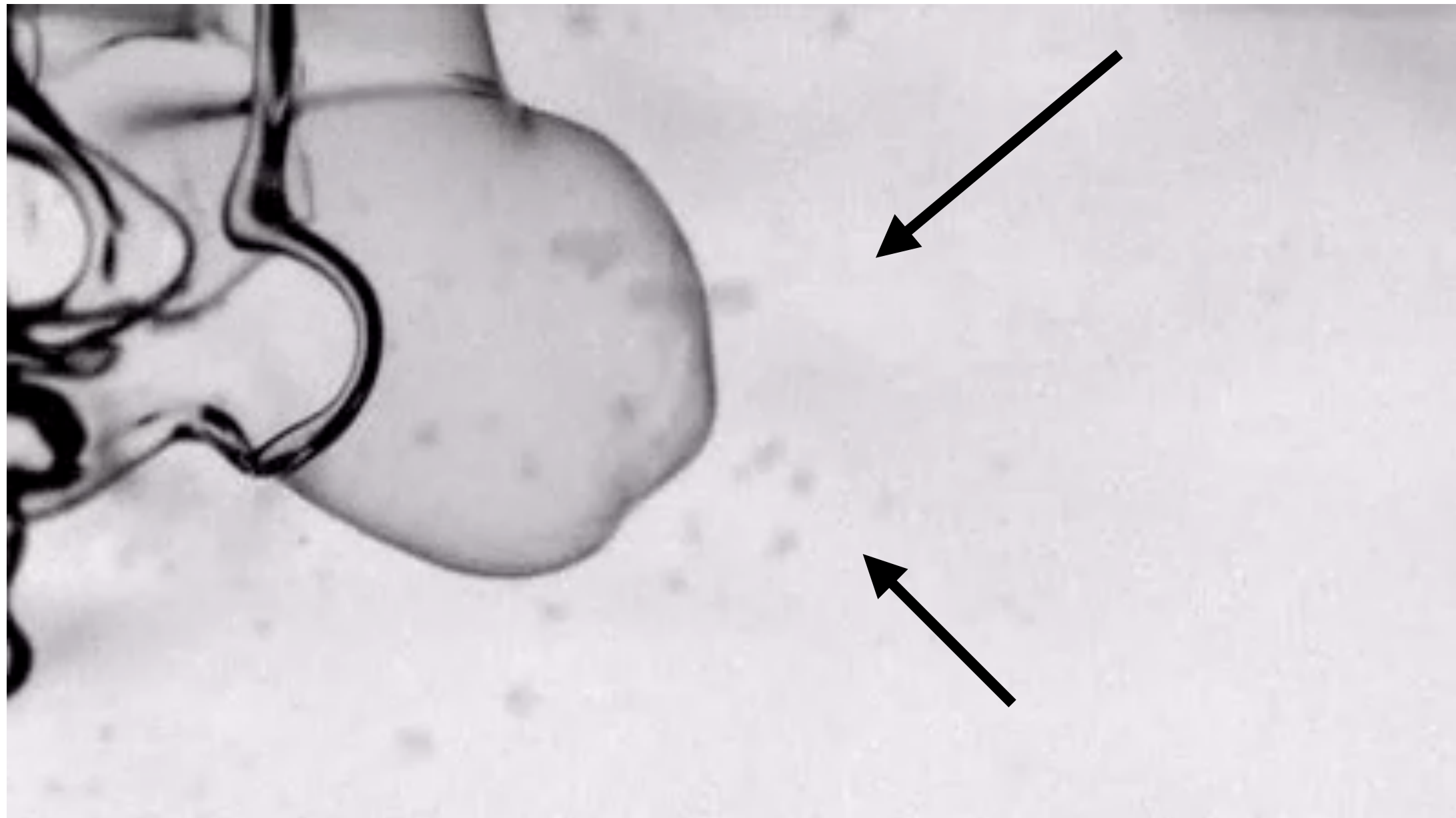


$$We = \frac{\rho_g R_0 U_0^2}{\gamma} \quad Oh = \frac{\eta_l}{\sqrt{\rho_l \gamma R_0}}$$

Hole nucleation in the bags



How do holes nucleate in sheets?



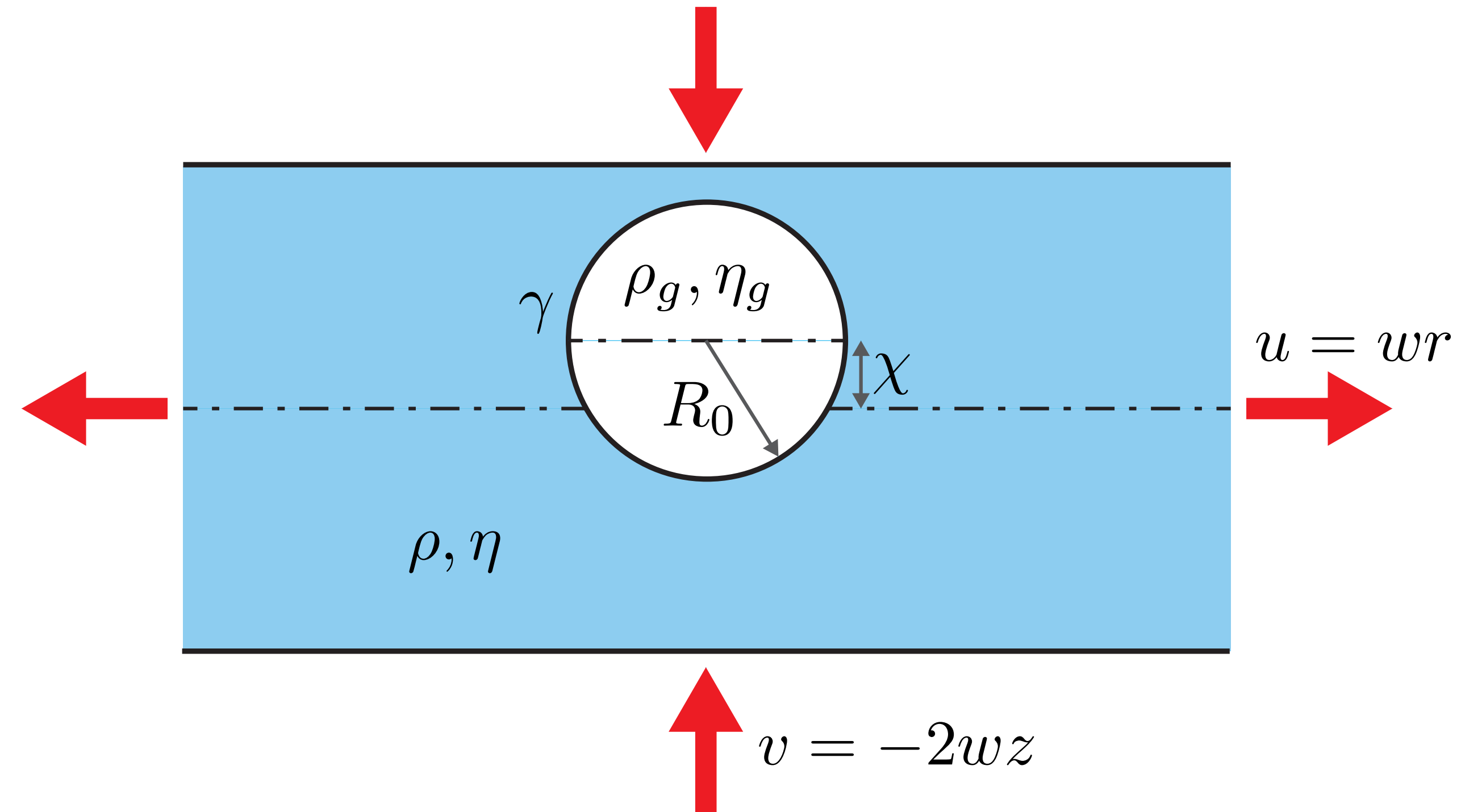
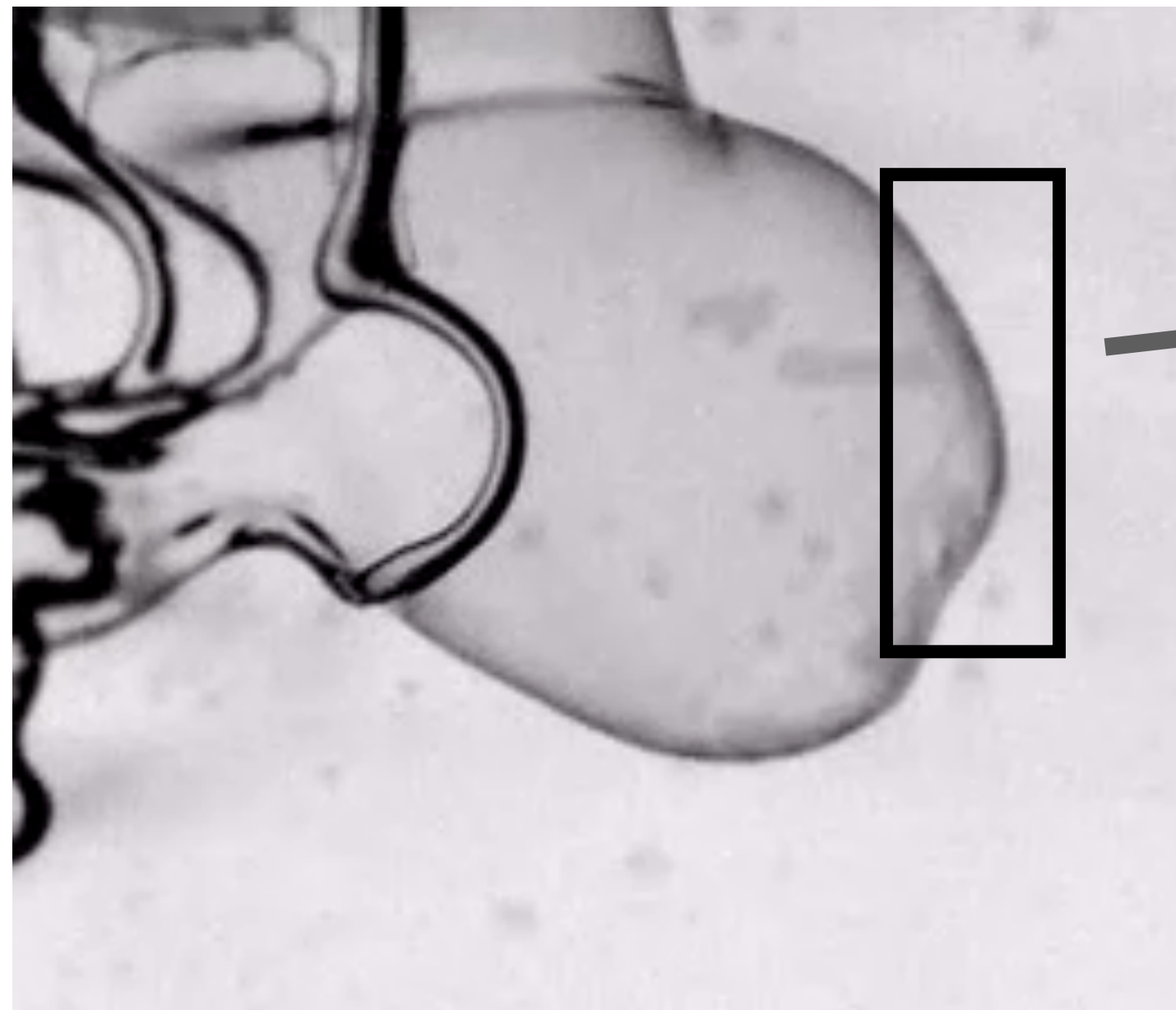
~~● Thermal fluctuations~~

~~● Van der Waal forces~~

● By hydrophobic particles or bubbles

● Chemical and temperature inhomogeneities

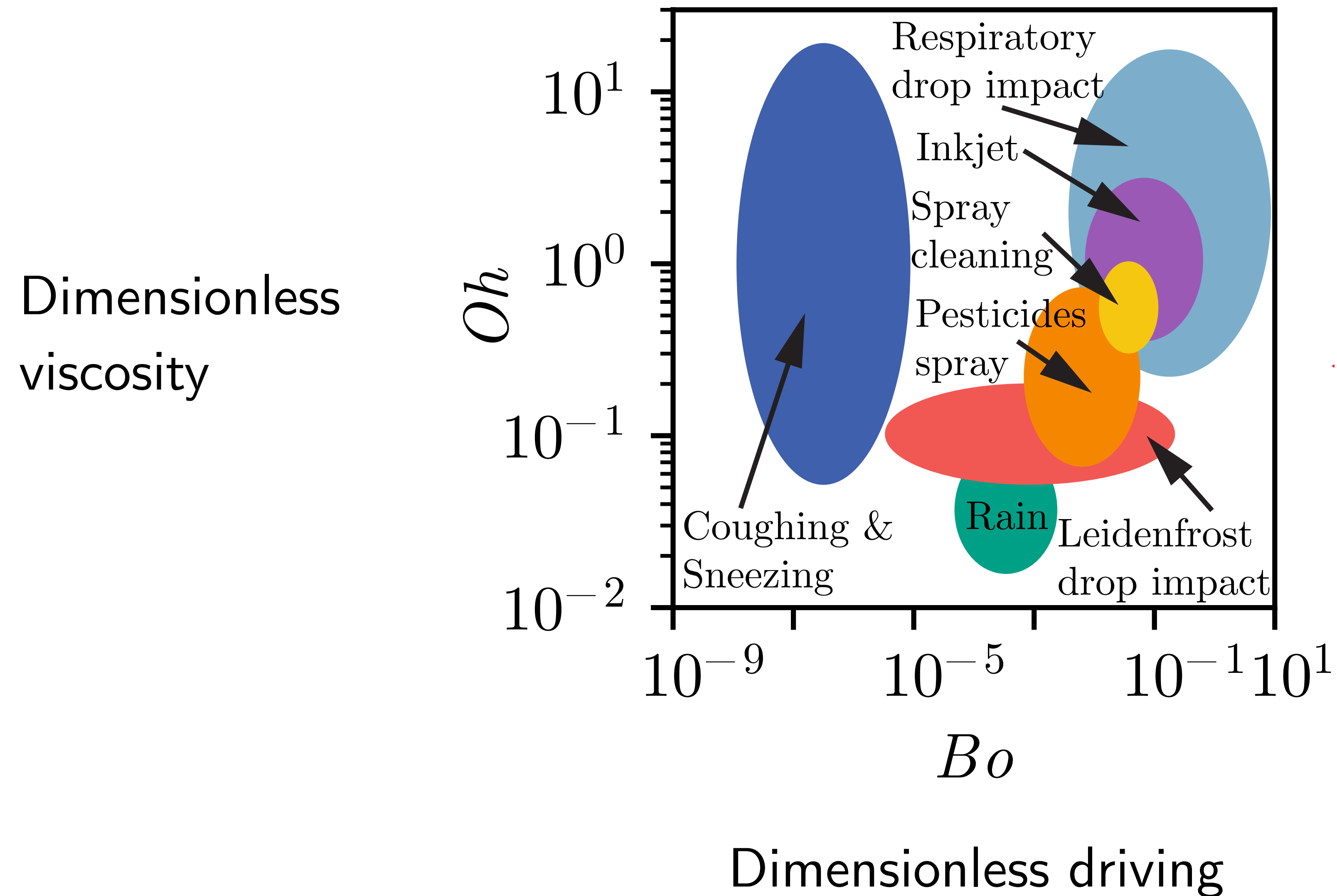
How to model this breakup?



$$Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}} \quad Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad \text{Asymmetry: } \frac{\chi}{R_0}$$

$$\frac{\rho_g}{\rho} = 10^{-3} \quad Oh_a = \frac{\eta_g}{\sqrt{\rho R_0 \gamma}} = 2 \times 10^{-5}$$

Relevance of omnipresent holey sheets

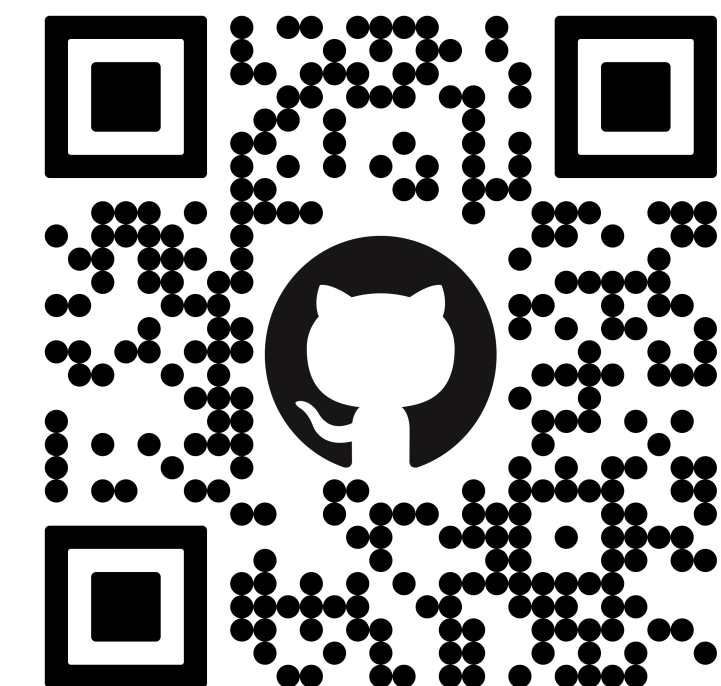


How do holey sheets break?

Computational Multiphase Physics Lab



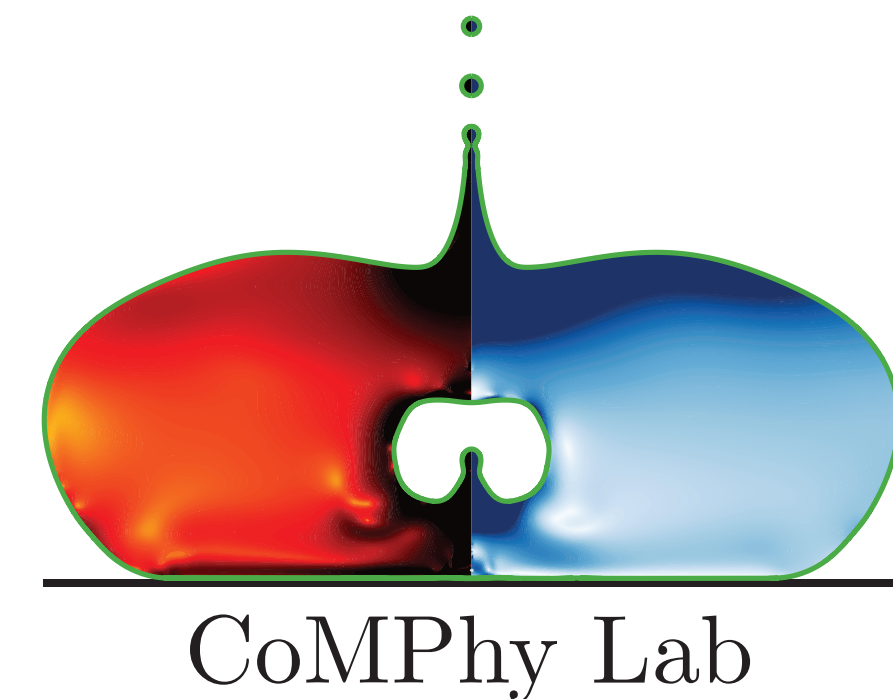
#ilovefs



Numerics: Basilisk C

Cauchy momentum + VoF

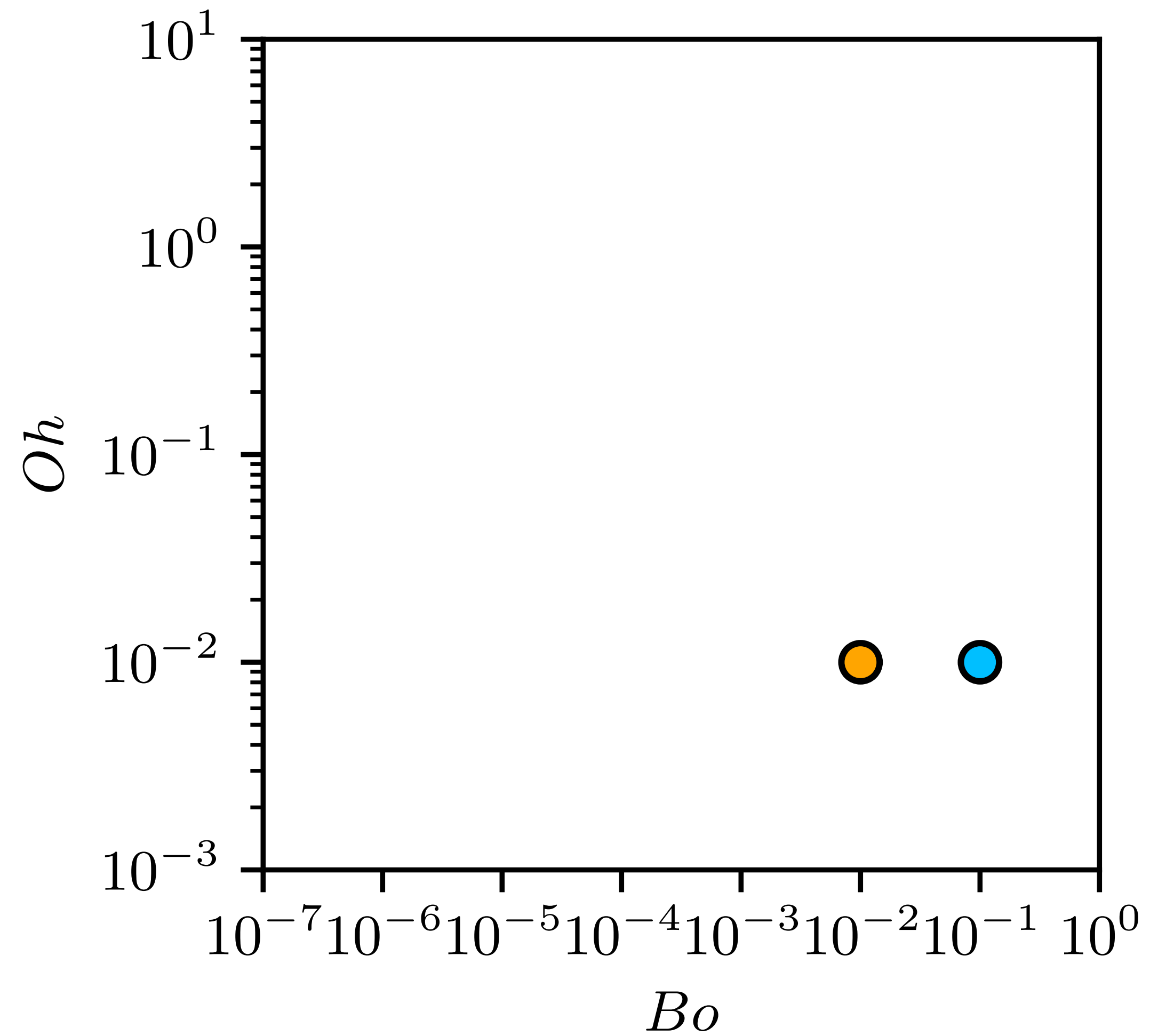
Stéphane Popinet & collaborators



CoMPhy Lab

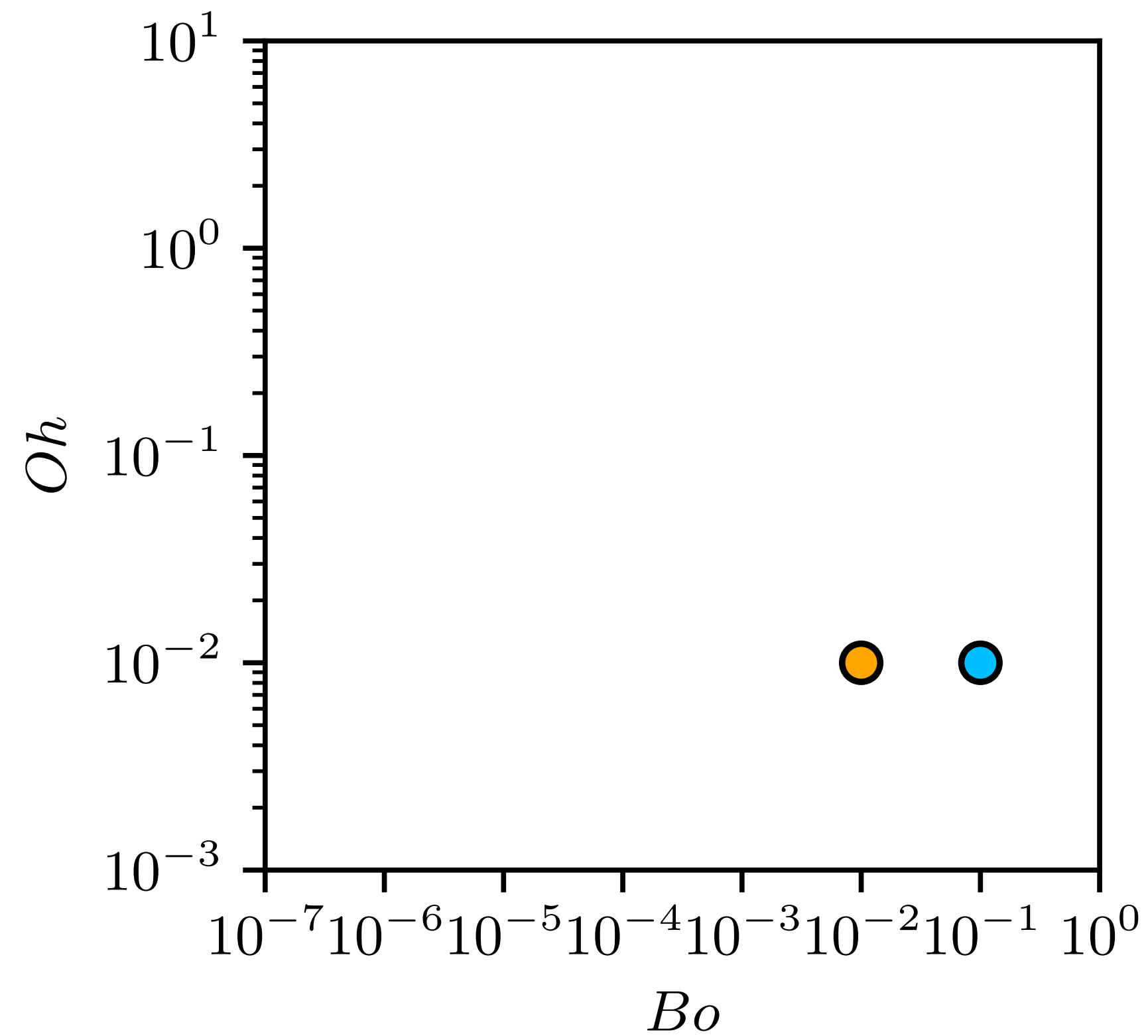
How do holey sheets break?

Oh-Bo parameter space

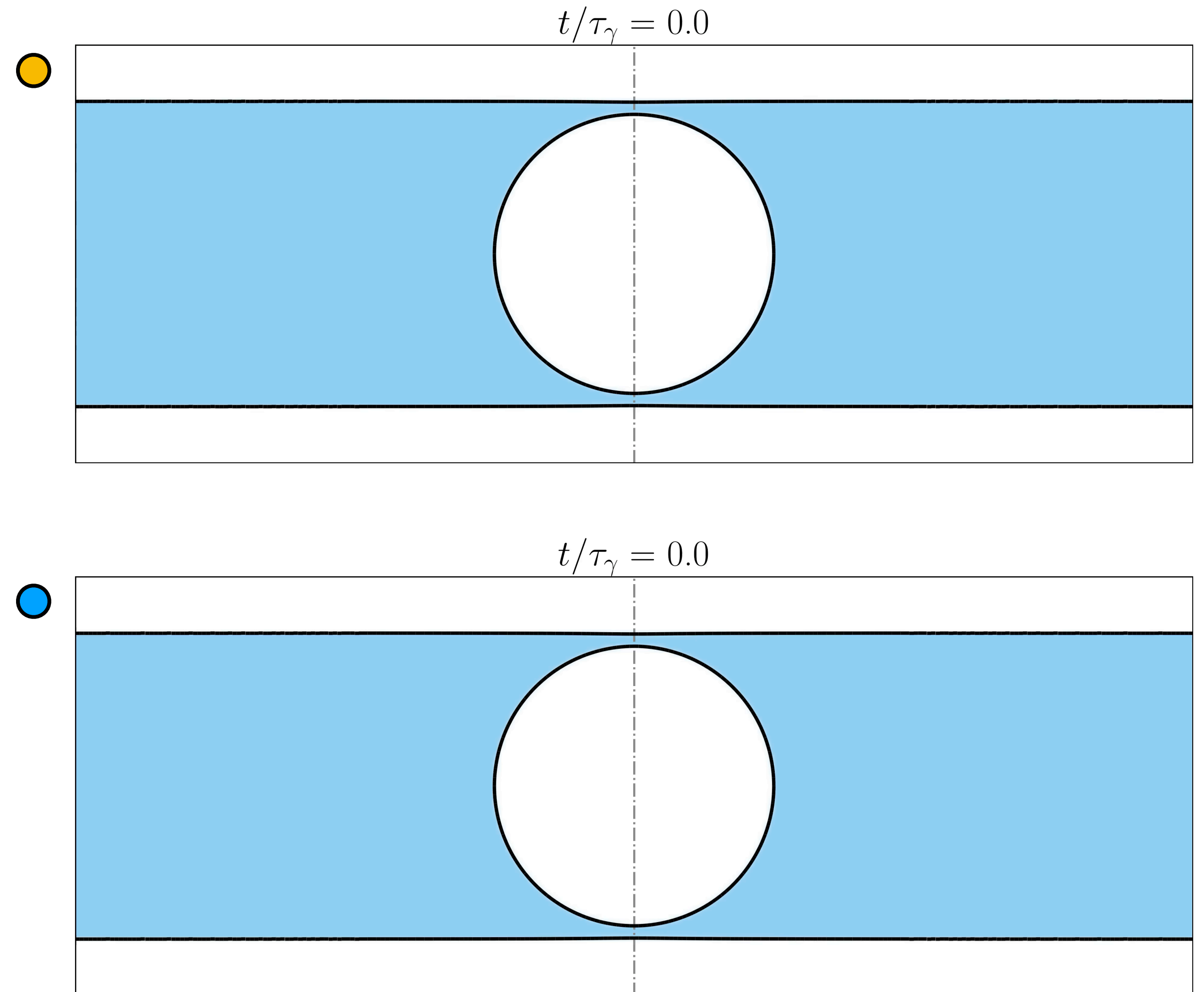


$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

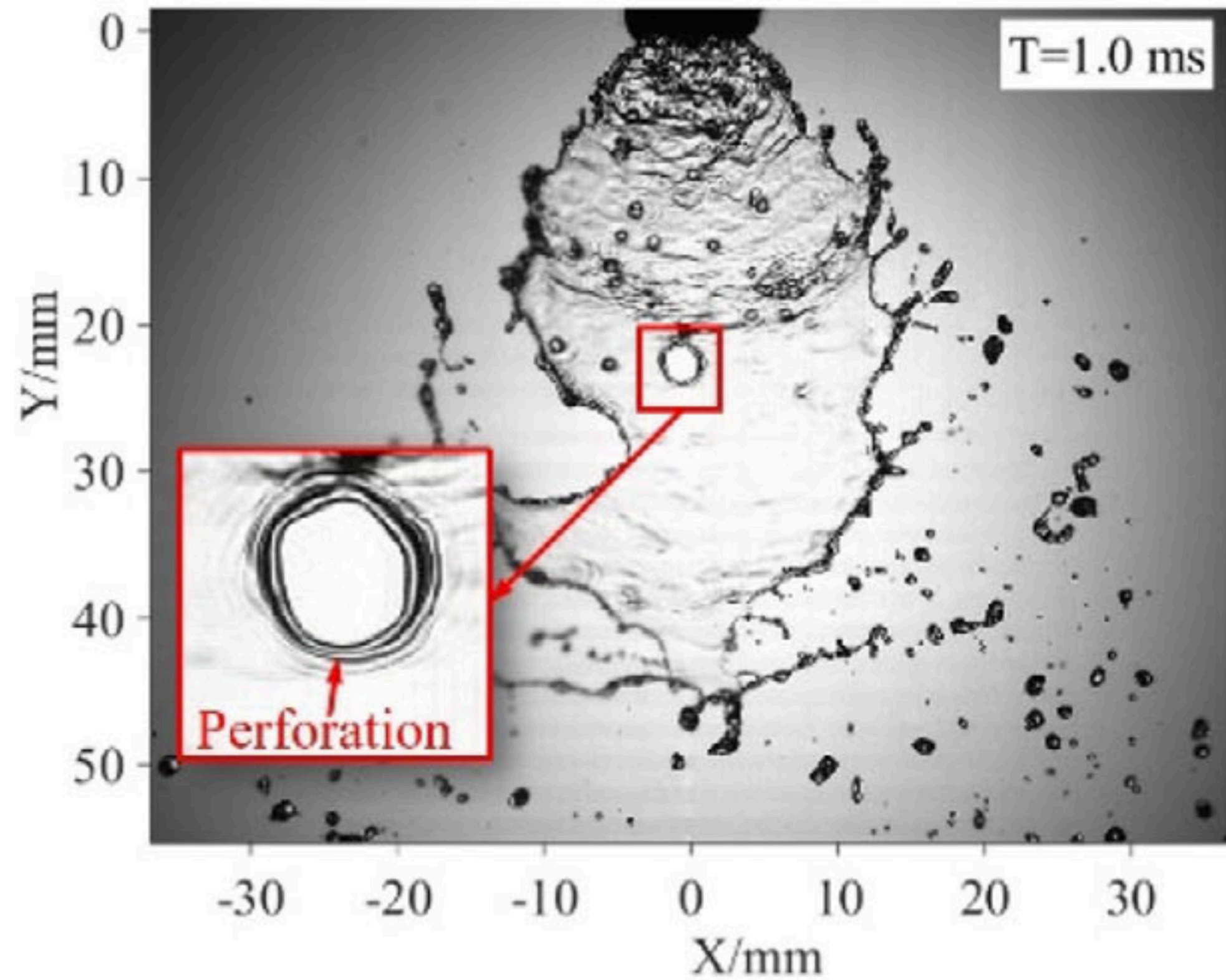
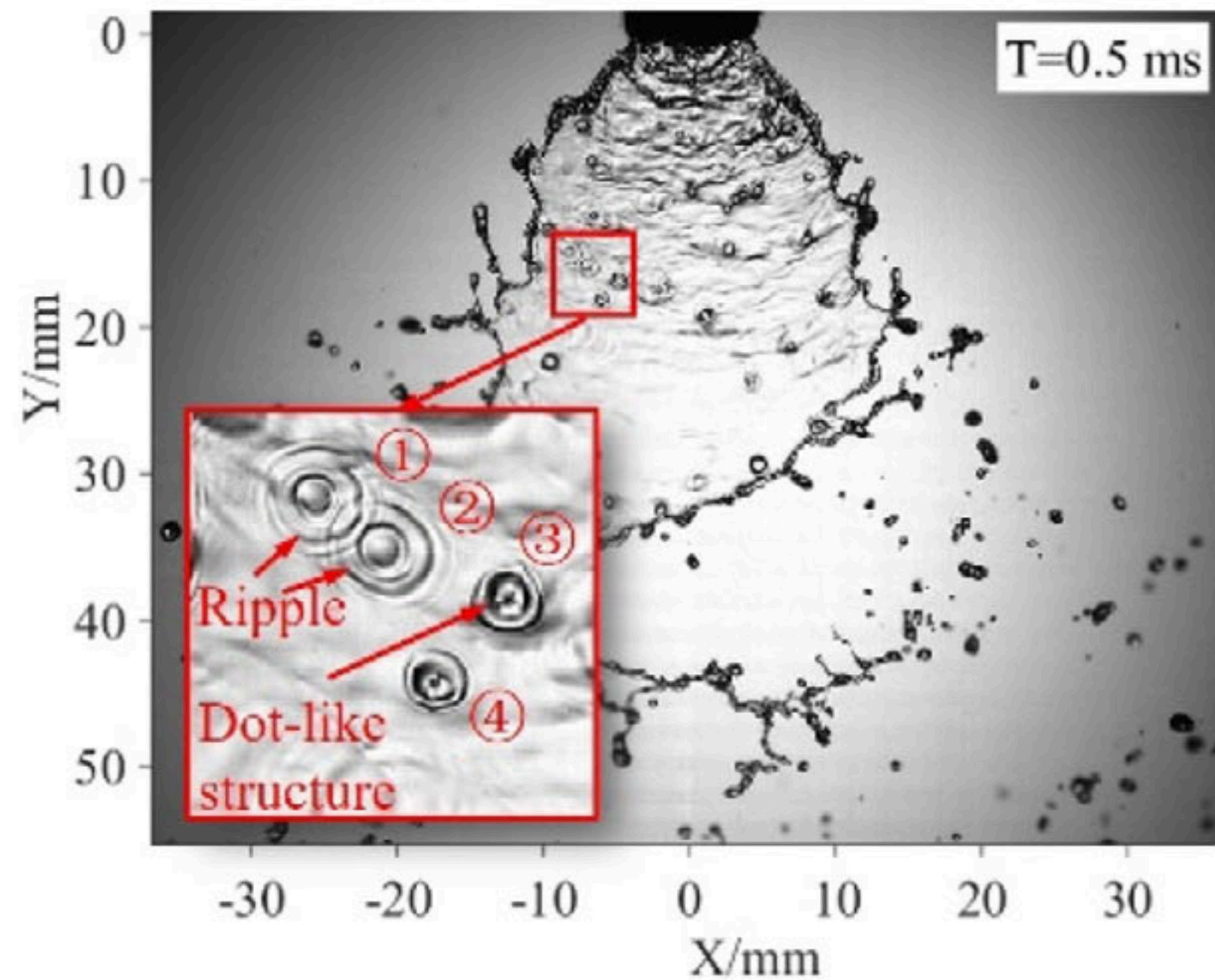
Oh - Bo parameter space



$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

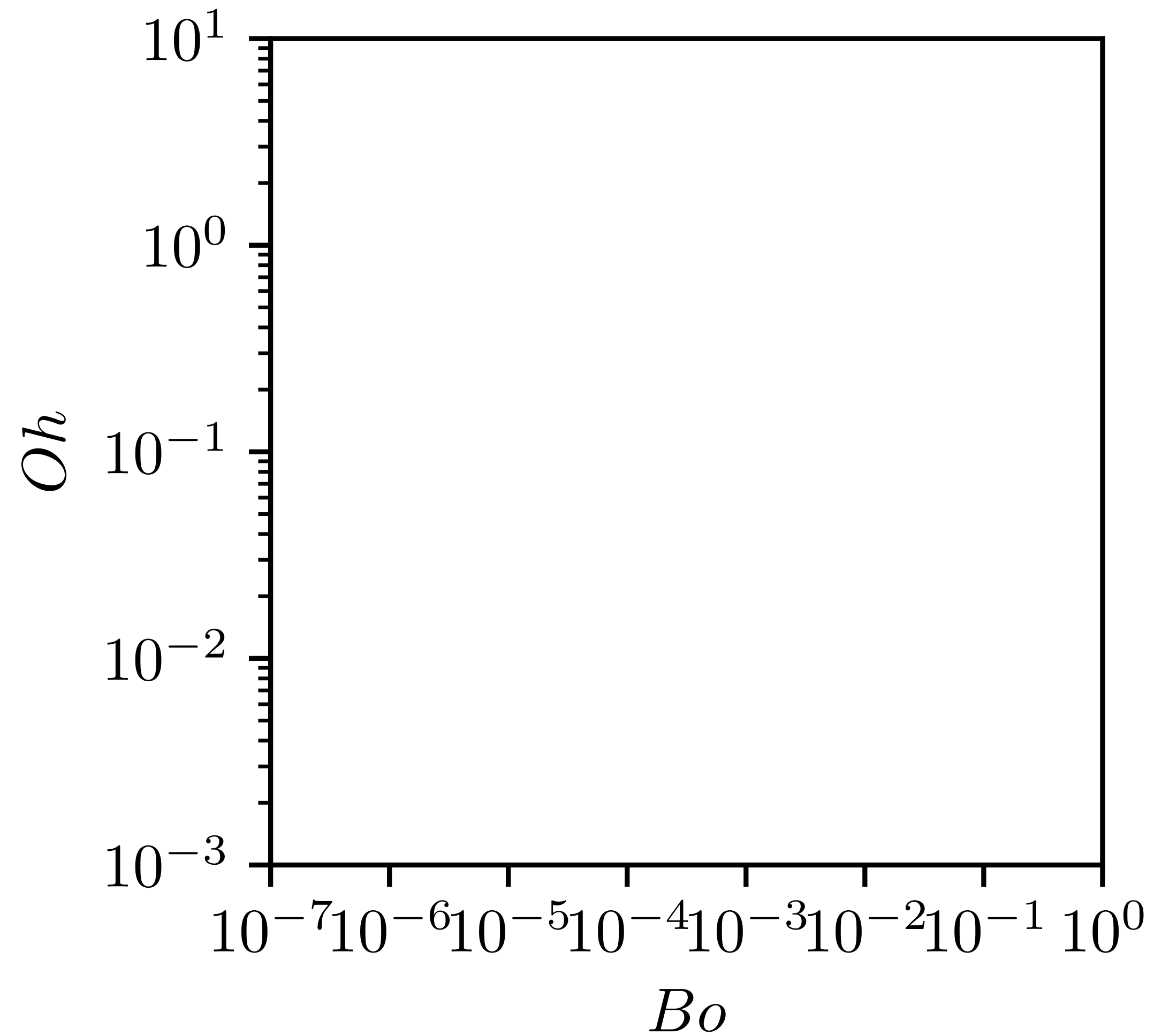


Experimental observations



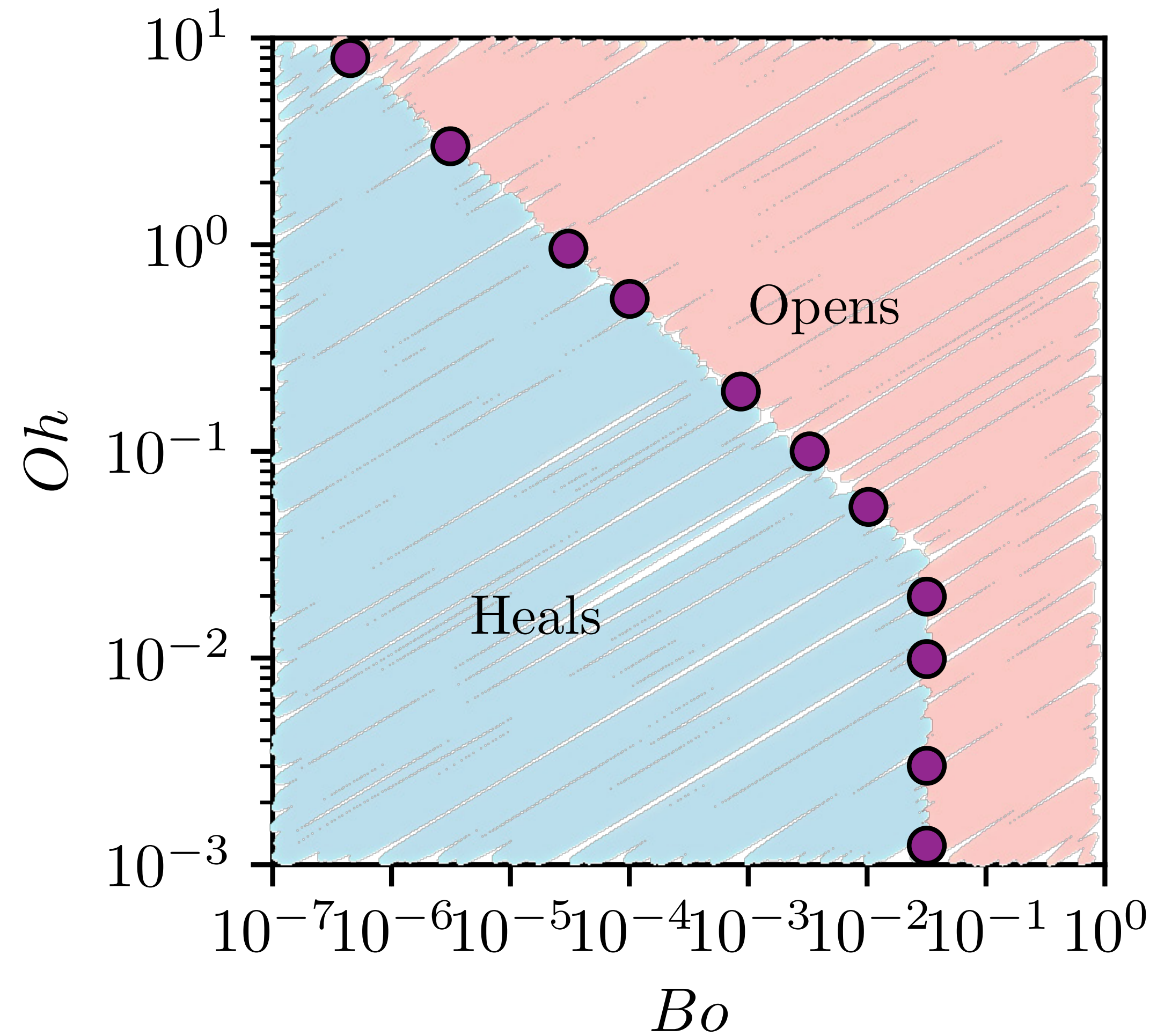
$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

Oh-Bo parameter space



$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

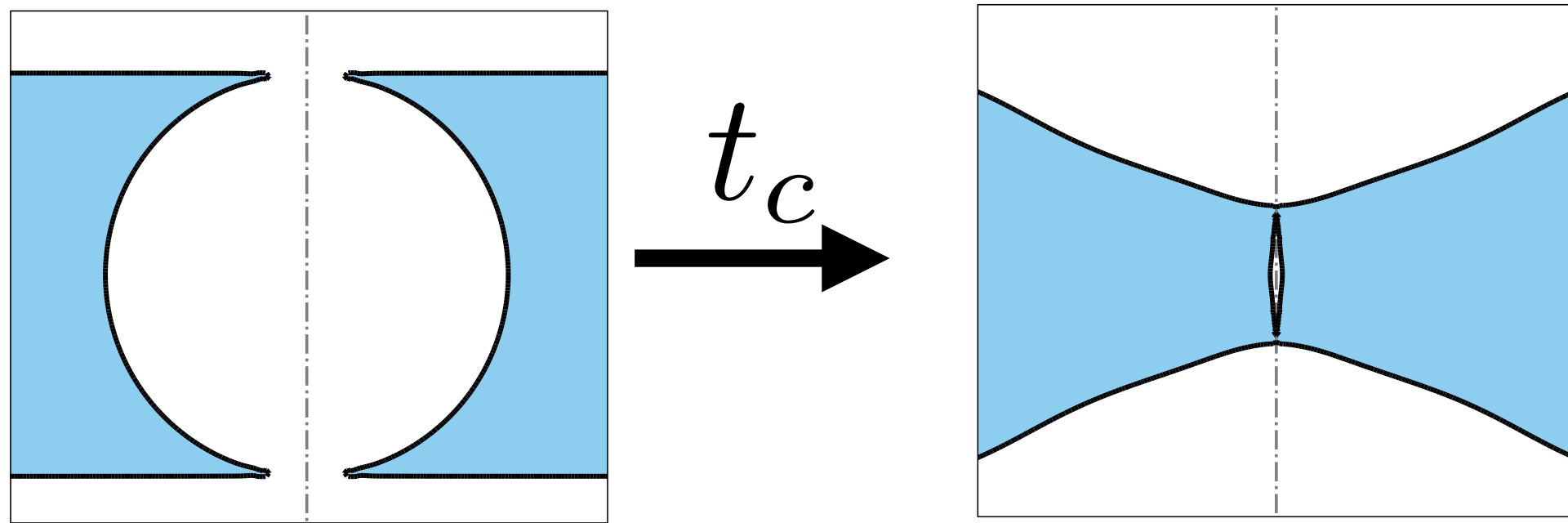
Oh - Bo parameter space: 300 simulations



$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

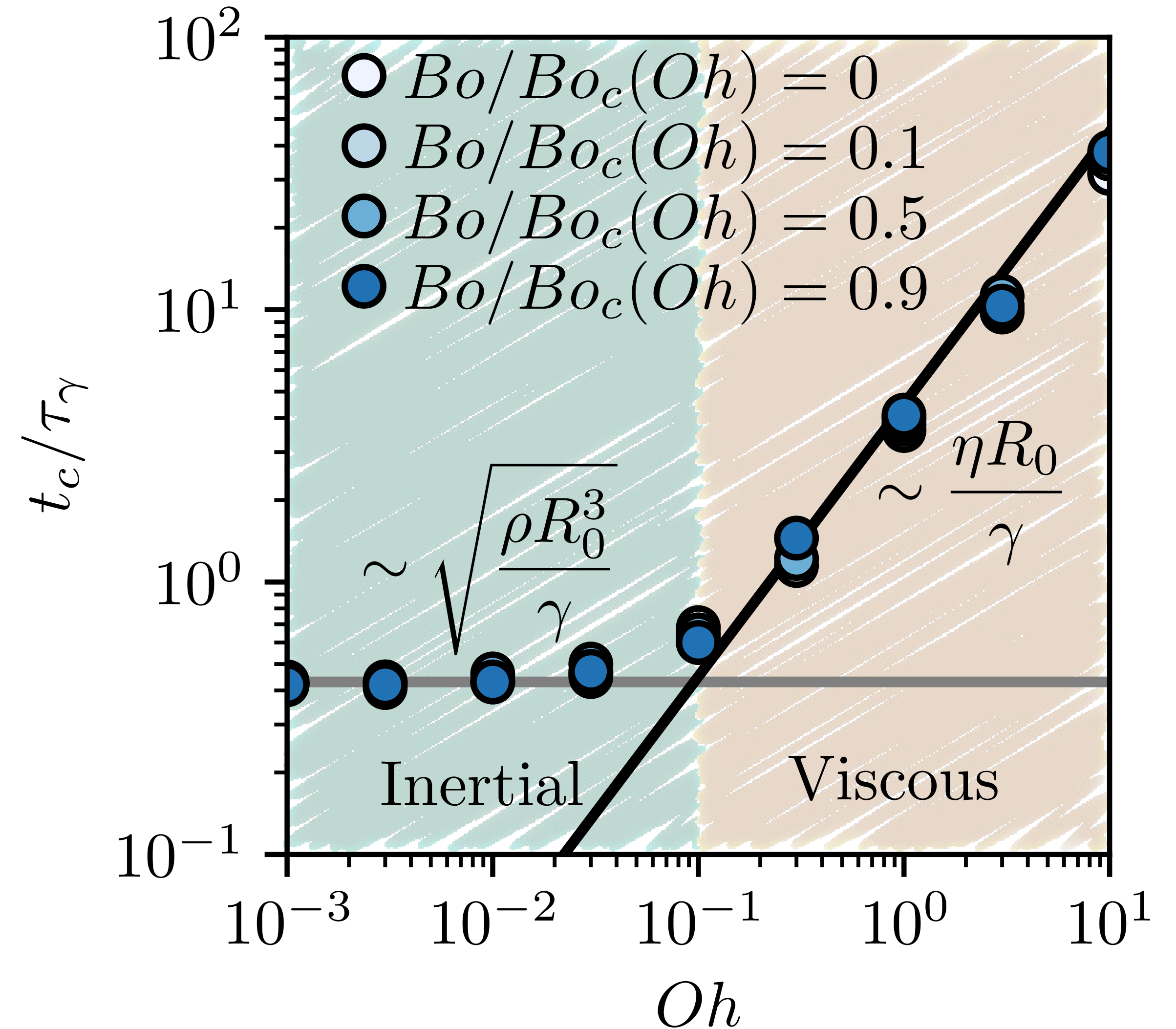
What sets the transitions?

Collision time scale

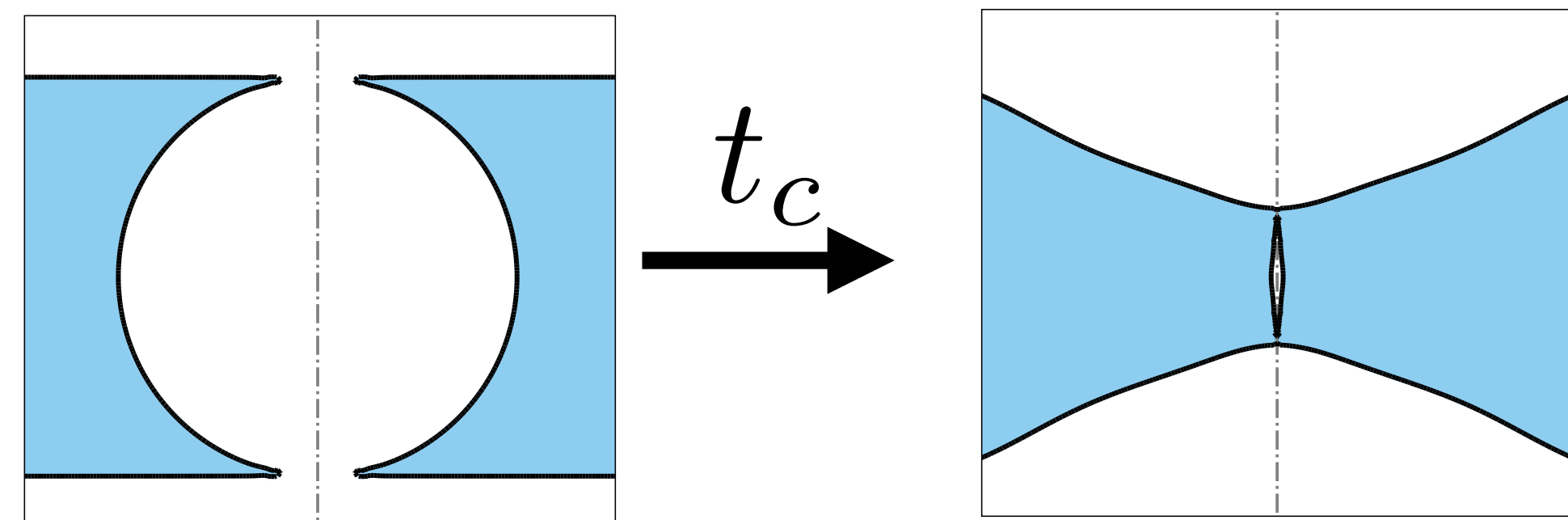


Small Oh : $\frac{\gamma}{R_0} \sim \rho U_0^2$ $t_c \sim \sqrt{\frac{\rho R_0^3}{\gamma}}$

Large Oh : $\frac{\gamma}{R_0} \sim \frac{\eta U_0}{R_0}$ $t_c \sim \frac{\eta R_0}{\gamma}$



Collision vs retraction at small Oh



Collision time scale $\sim$ Retraction time scale of tip

$$\sqrt{\frac{\rho R_0^3}{\gamma}} \sim \frac{1}{\omega}$$

$$\frac{\rho \omega^2 R_0^3}{\gamma} = \text{constant}$$

$$Bo = \text{constant}$$

$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

Collision vs retraction at high Oh

Collision time scale $\sim$ Retraction time scale of tip

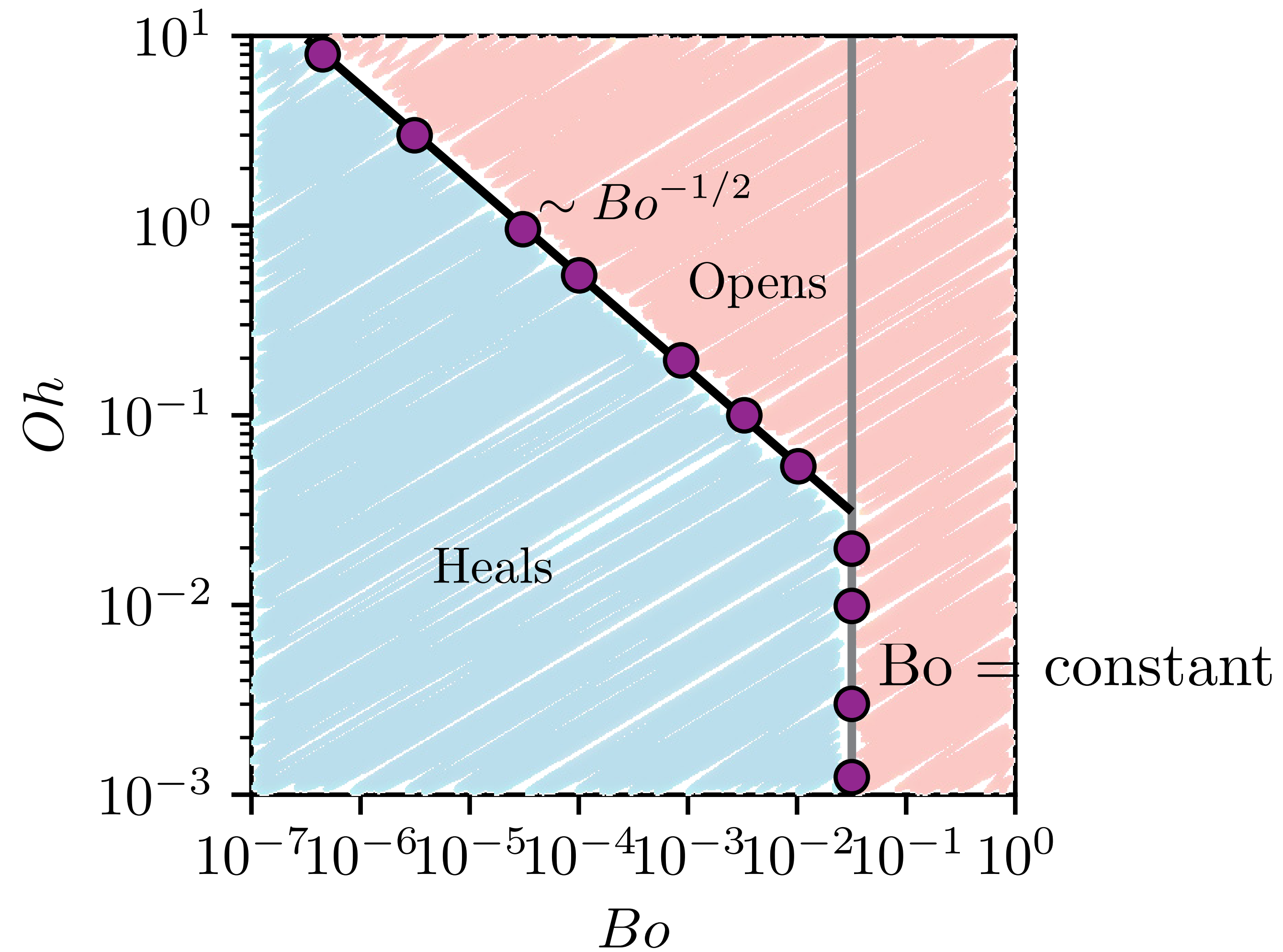
$$\frac{\eta R_0}{\gamma} \sim \frac{1}{\omega}$$

$$\frac{\eta R_0 \omega}{\gamma} \sim 1 \quad Oh \sim Bo^{-1/2}$$

$$Oh Bo^{1/2} = \text{constant}$$

$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

Oh-Bo parameter space

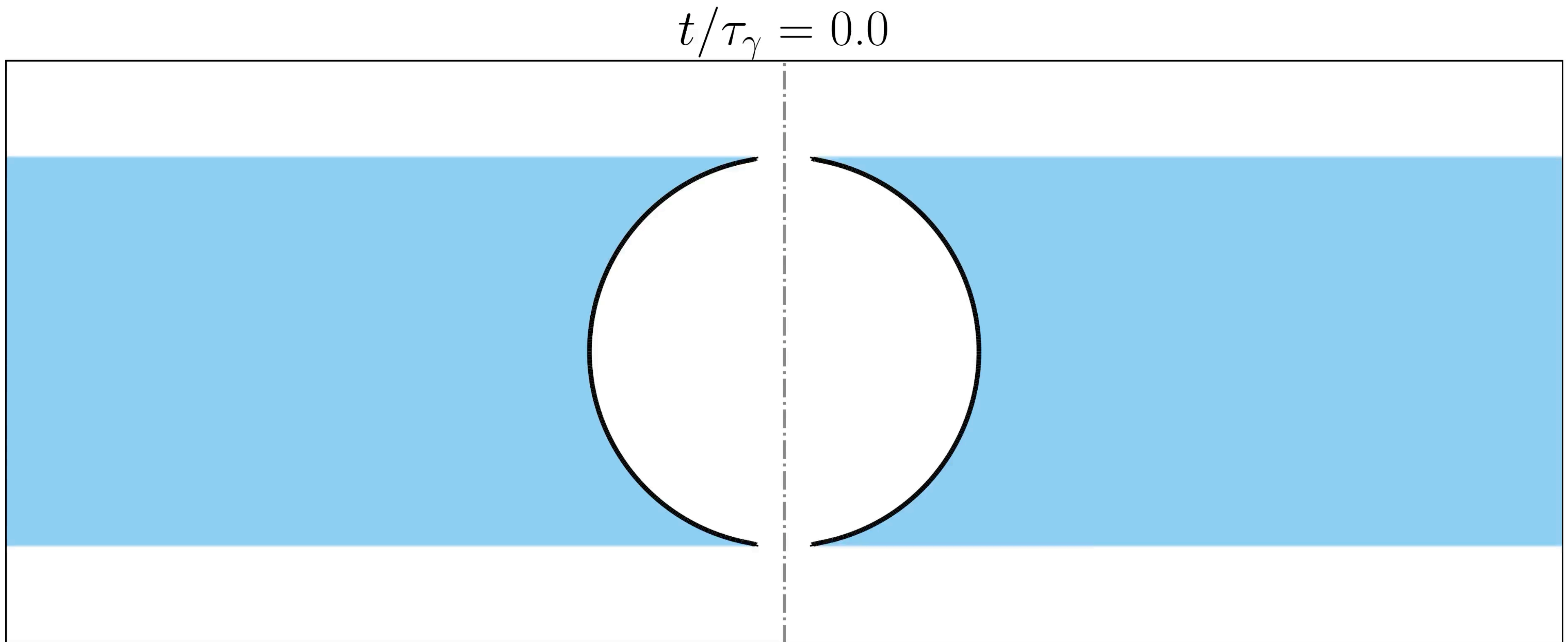


$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

$$Oh_c \sim 1/\sqrt{Bo}$$

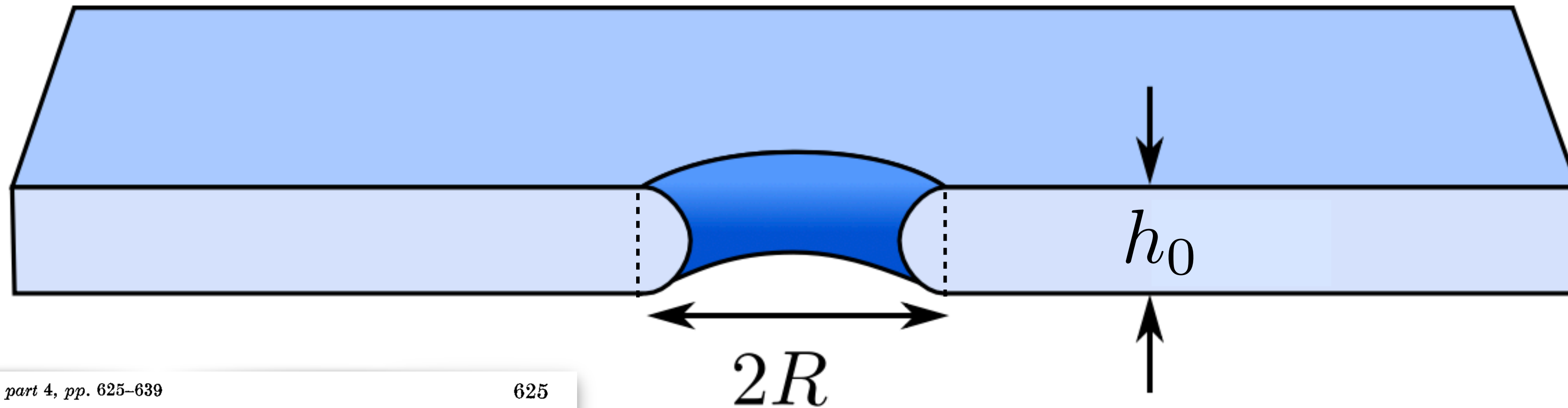
What happens for vanishing driving?

Holes always heal at $Bo = 0$



$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

Taylor & Michael's geometric condition for hole opening



J. Fluid Mech. (1973), vol. 58, part 4, pp. 625–639
Printed in Great Britain

625

On making holes in a sheet of fluid

By G. I. TAYLOR

Trinity College, Cambridge

AND D. H. MICHAEL

Department of Mathematics, University College London

(Received 12 October 1972)

It is suggested in this paper that axisymmetric holes in thin sheets of fluid in which surface tension forces predominate will open out if they are initially large in relation to the thickness of the sheet; but that small holes will close up. No exact criterion has been found for the critical hole size in a free falling sheet, but the behaviour of the sheet may be closely simulated by the suspension of a soap film between coaxial circular rings. Theoretical results and experimental observations on catenoid films so formed are described.

$$\frac{\Delta G}{\gamma h_0^2} = -\pi + \pi^2 \left(\frac{R}{h_0} \right) - 2\pi \left(\frac{R}{h_0} \right)^2$$

$$\frac{R}{h_0} \gg \frac{\pi}{4} - \frac{1}{2}$$

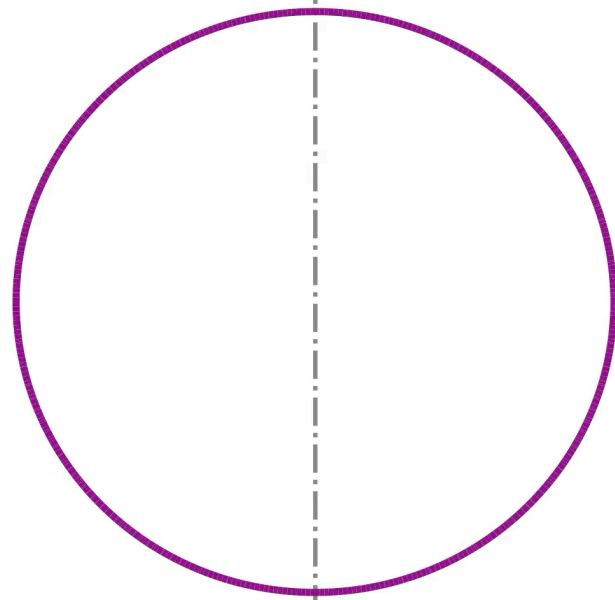
Not included: dynamic effects

What if the hole geometry changes?

If there is delayed breakup!

$$Oh = 0.1 \quad Bo = 0.1$$

$$t/\tau_\gamma = 0.100$$



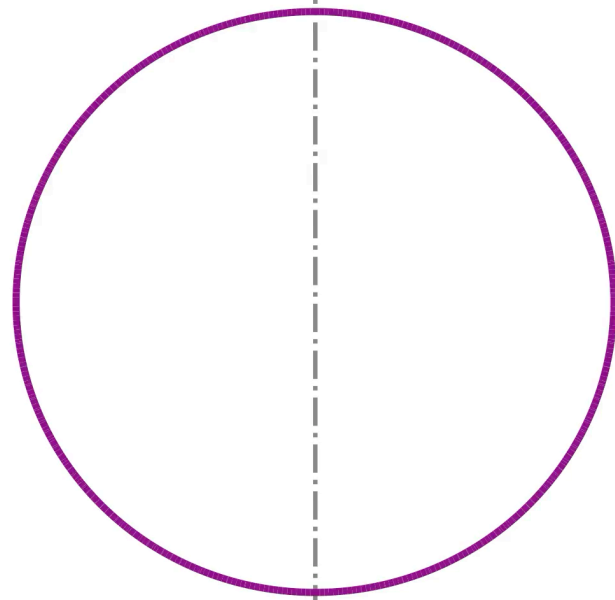
$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

Time-controlled rupture

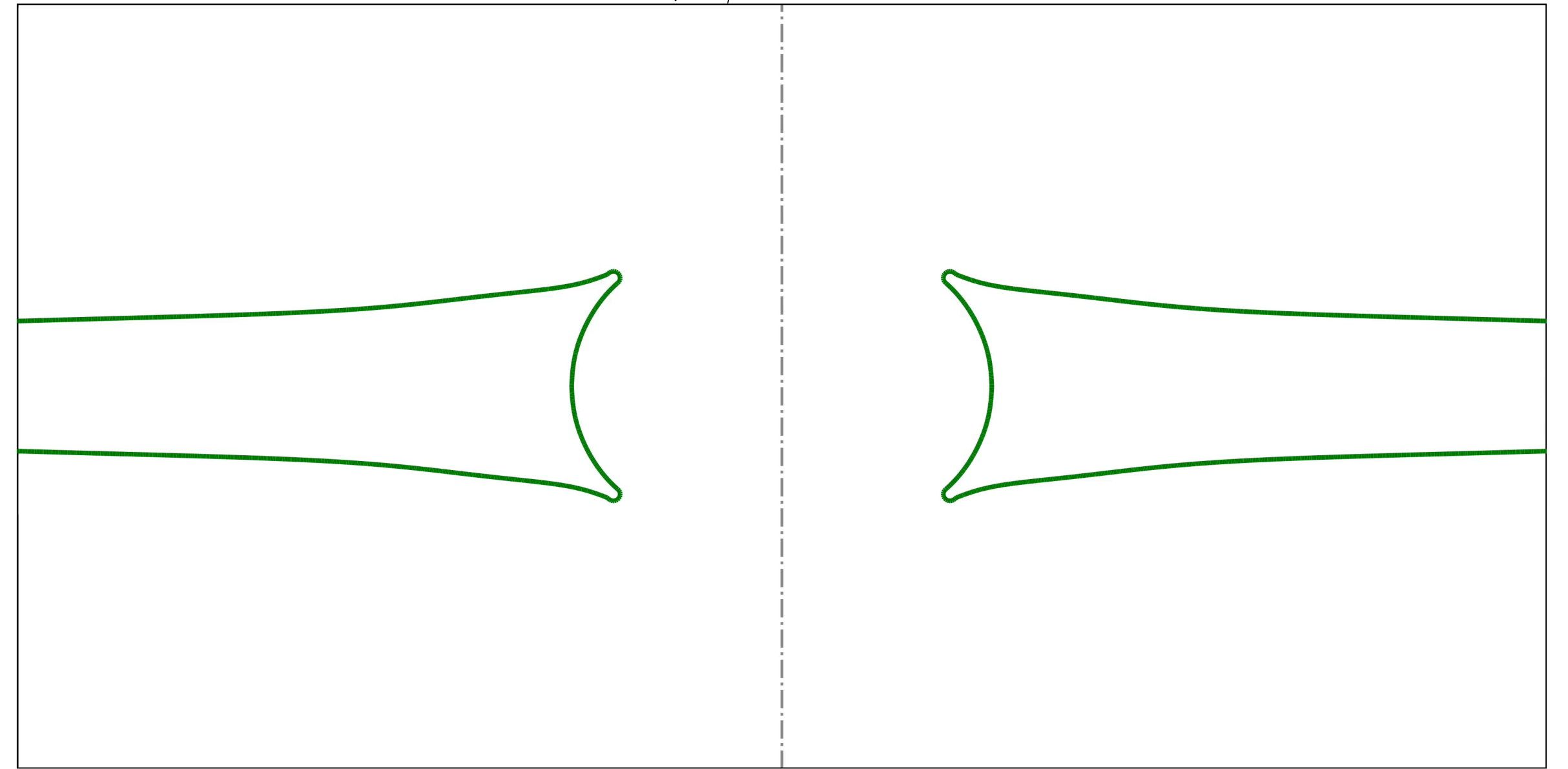
$$Oh = 0.1 \quad Bo = 0.1$$

$$t_r/\tau_\gamma = 6$$

$$t/\tau_\gamma = 0.100$$

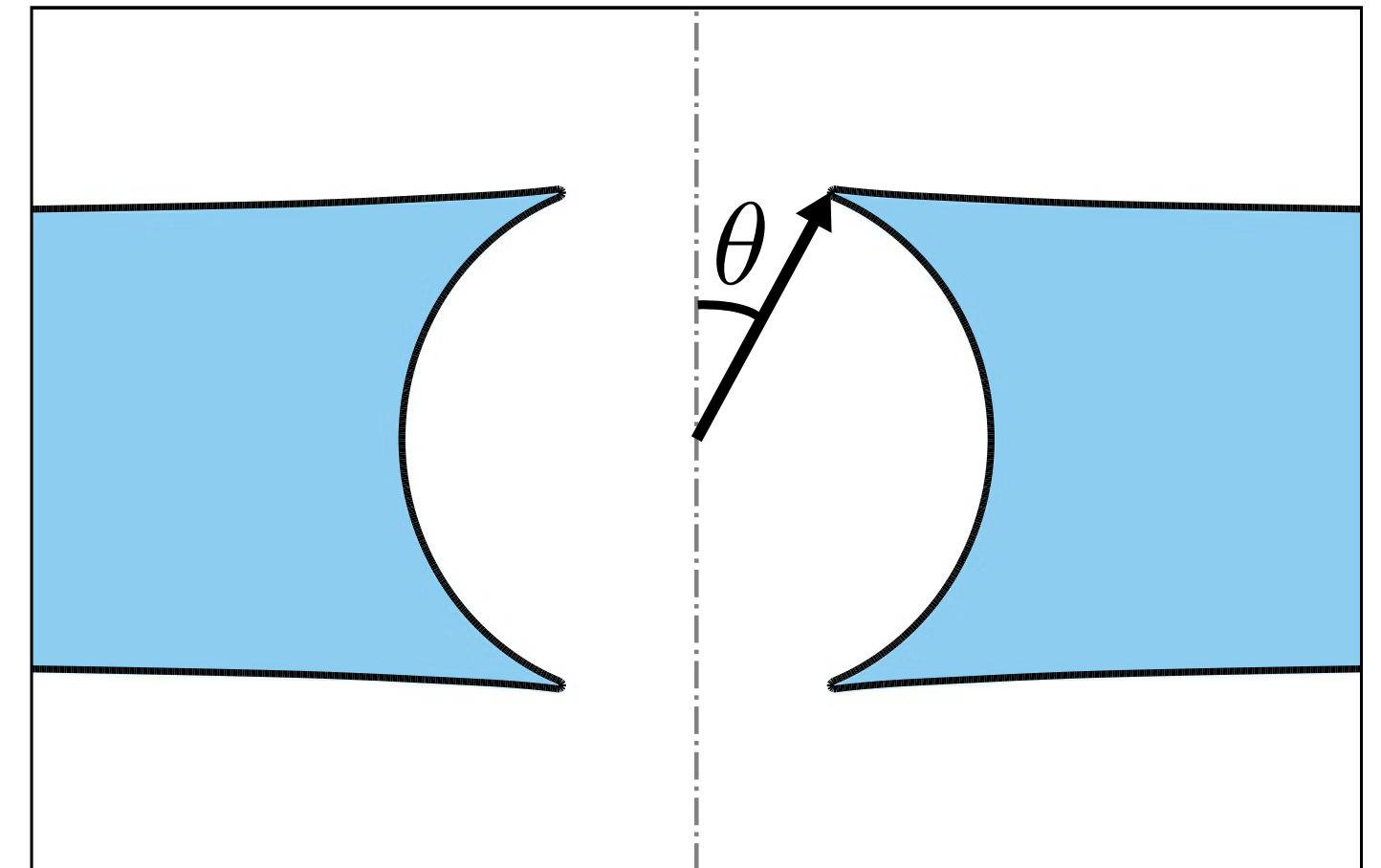
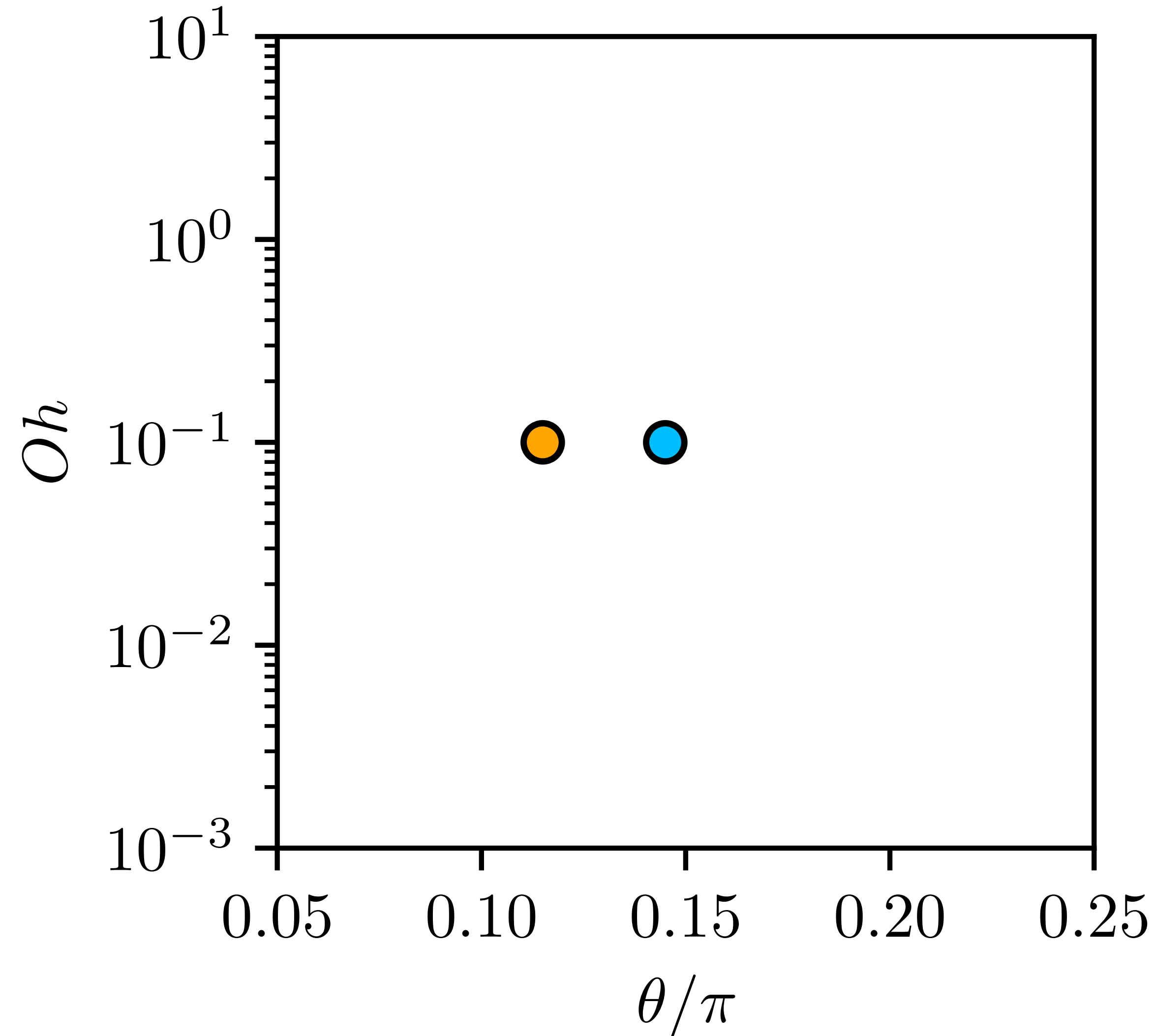


$$t/\tau_\gamma = 0.000$$



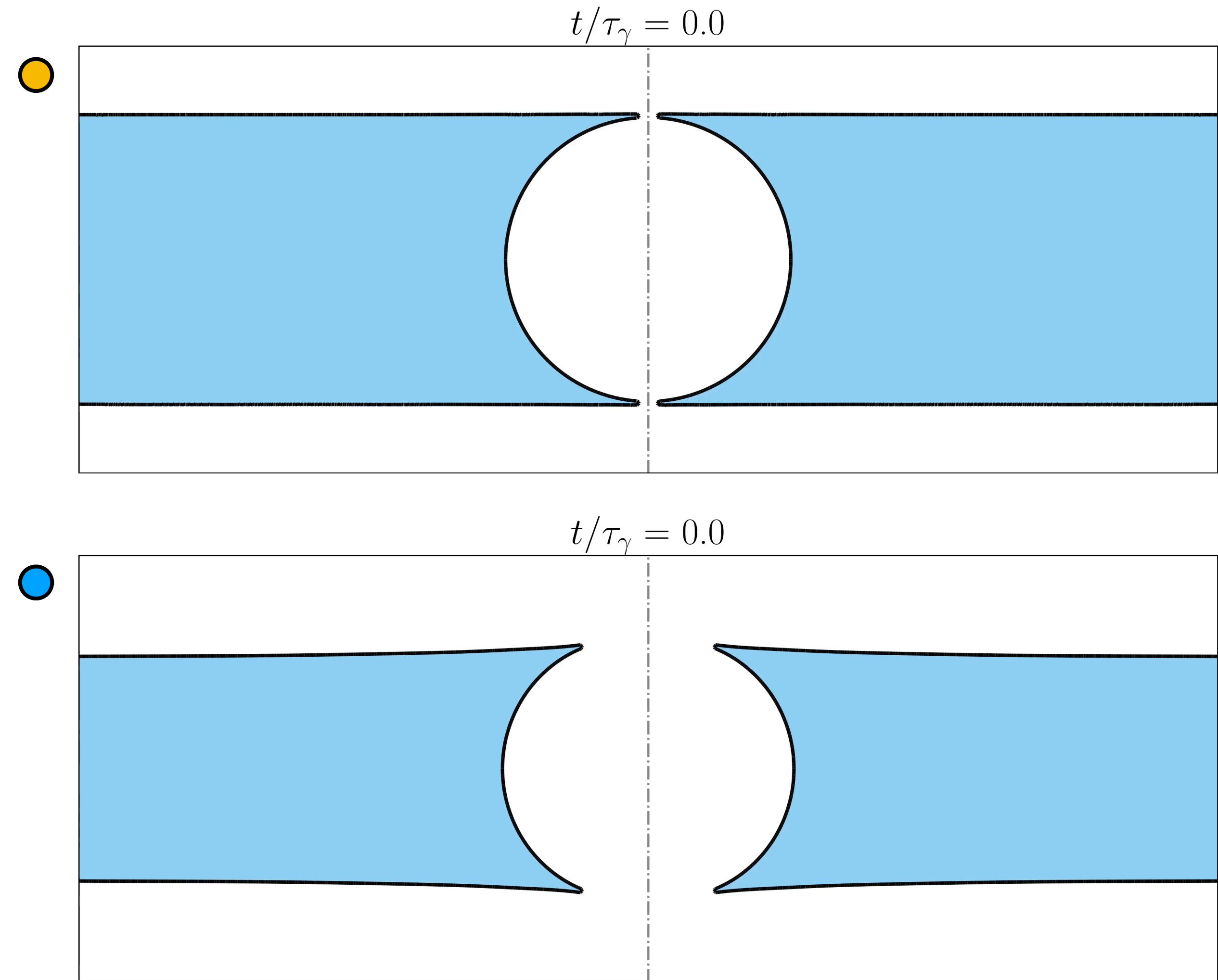
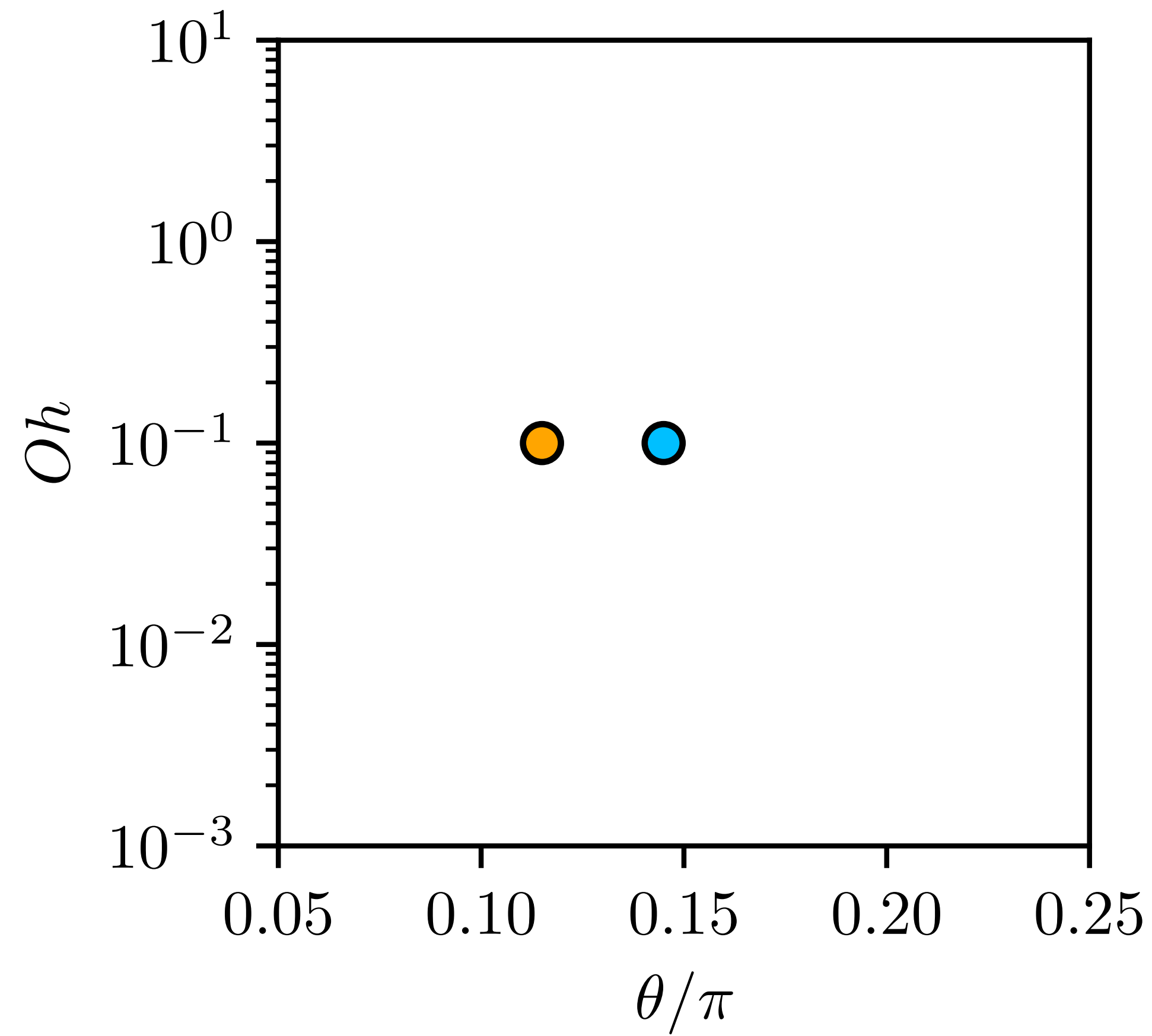
$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

Initial shape effects at $Bo = 0$



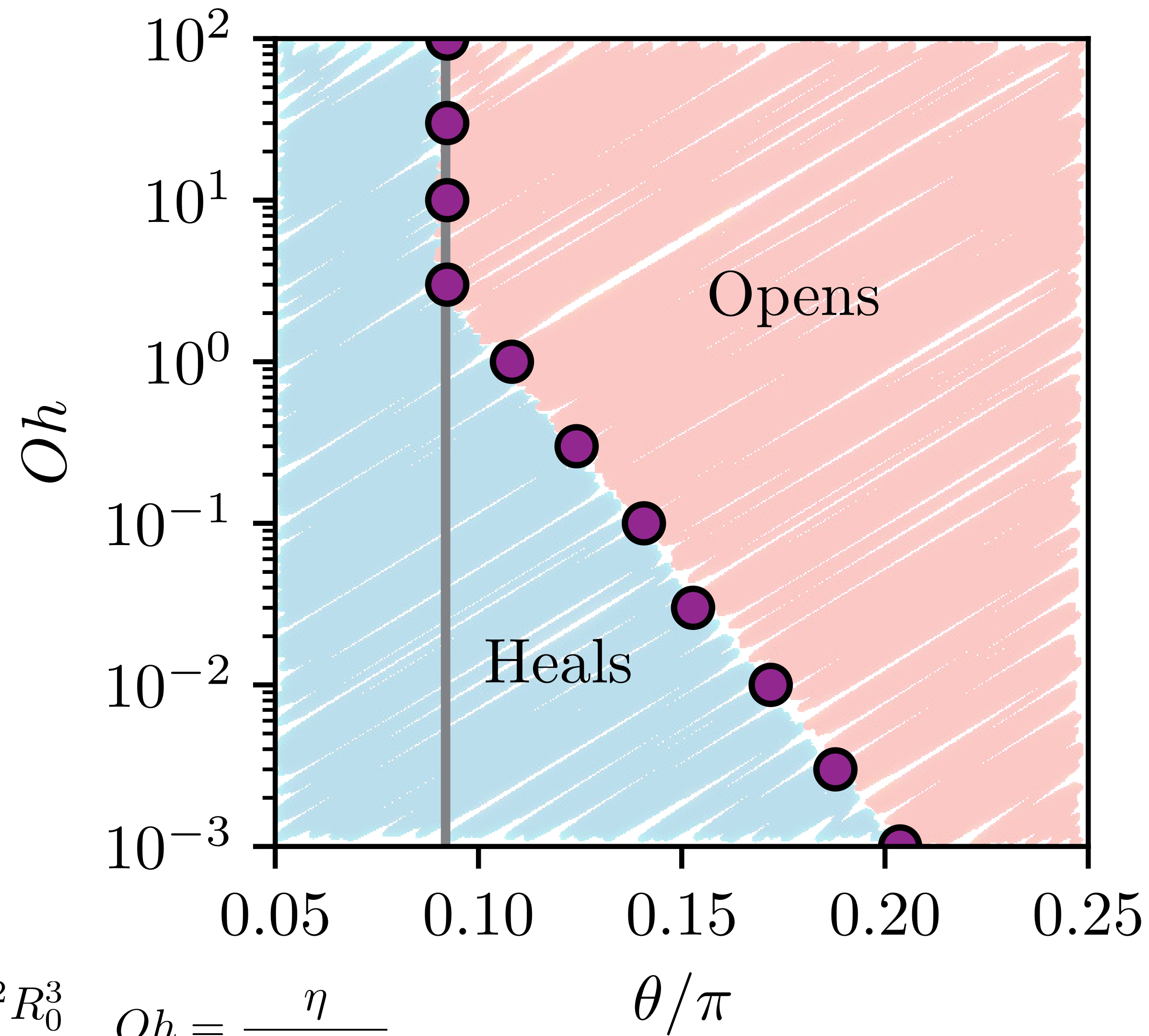
$$Bo = \frac{\rho \omega^2 R_0^2}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

Initial shape effects at $Bo = 0$

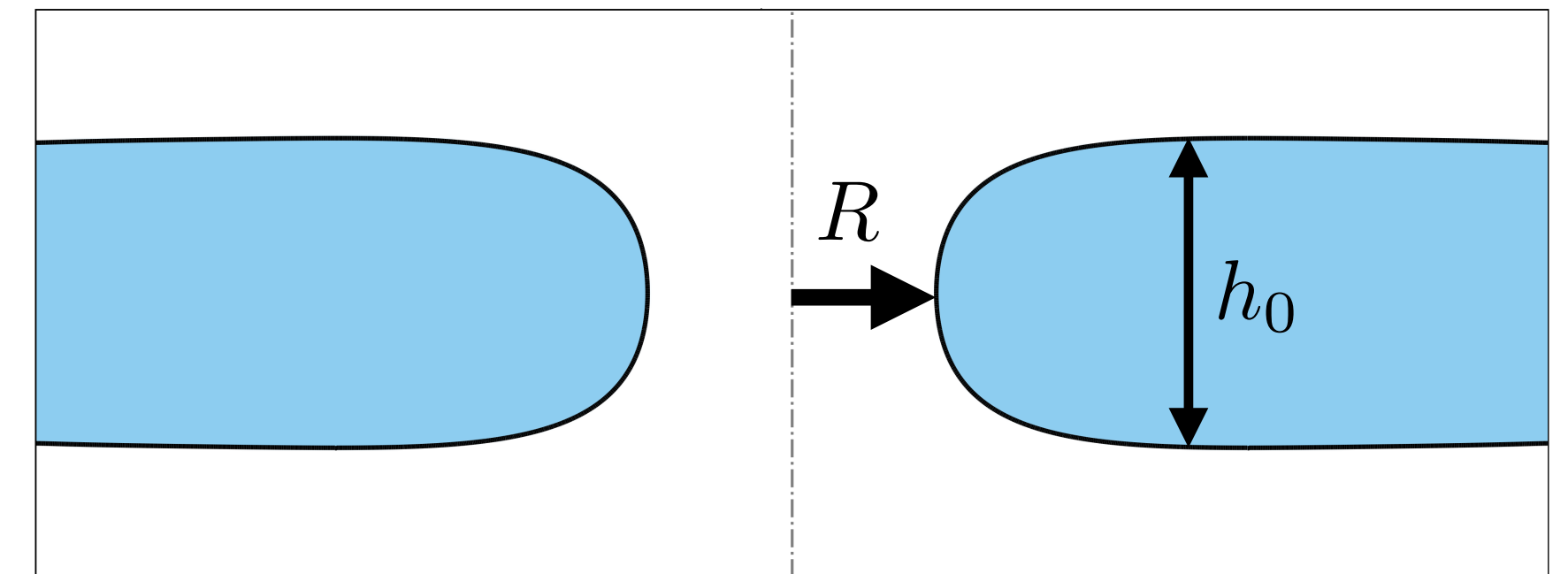


$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

Regime map (Oh - Bo)



Near transition high Oh cases



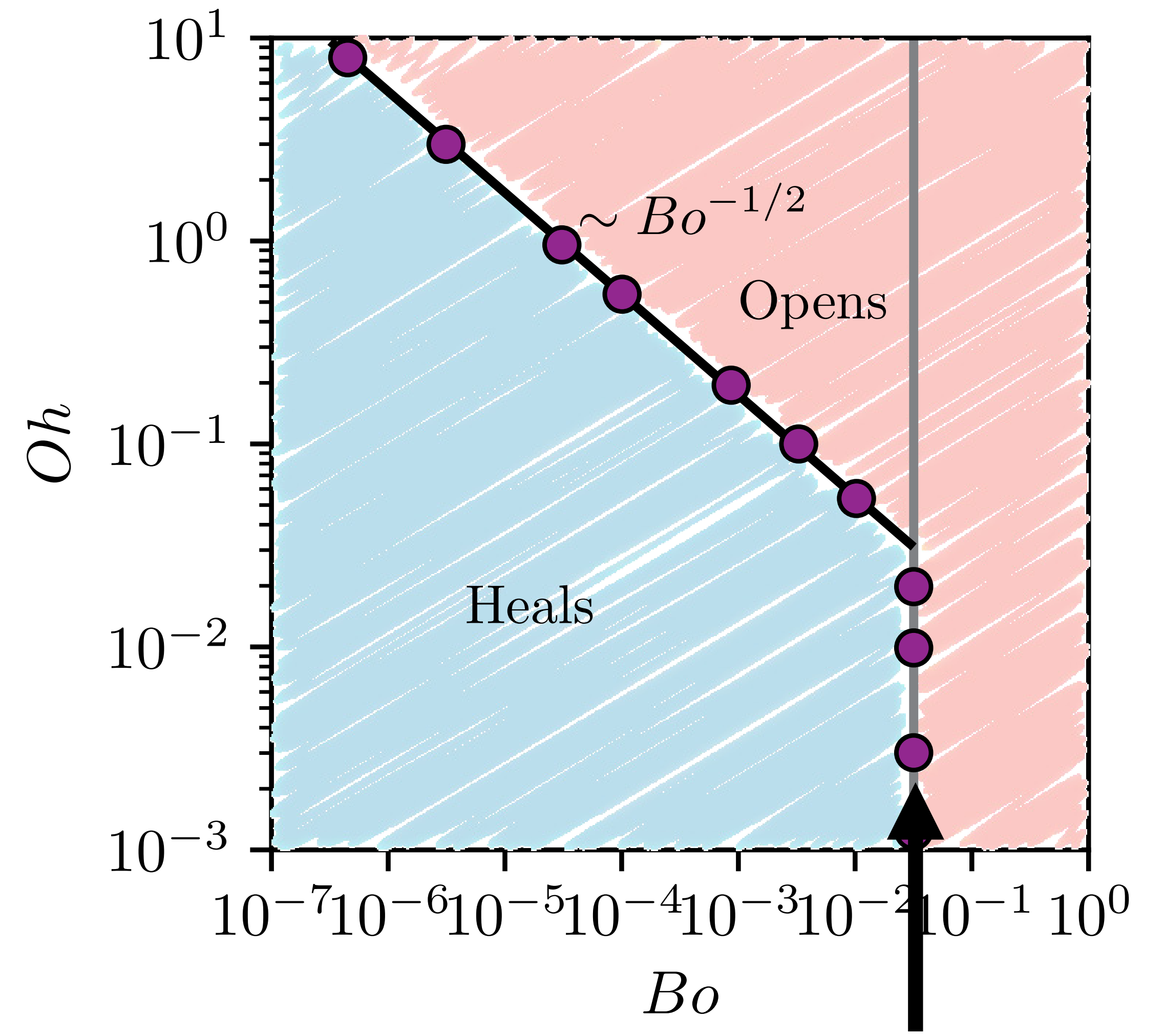
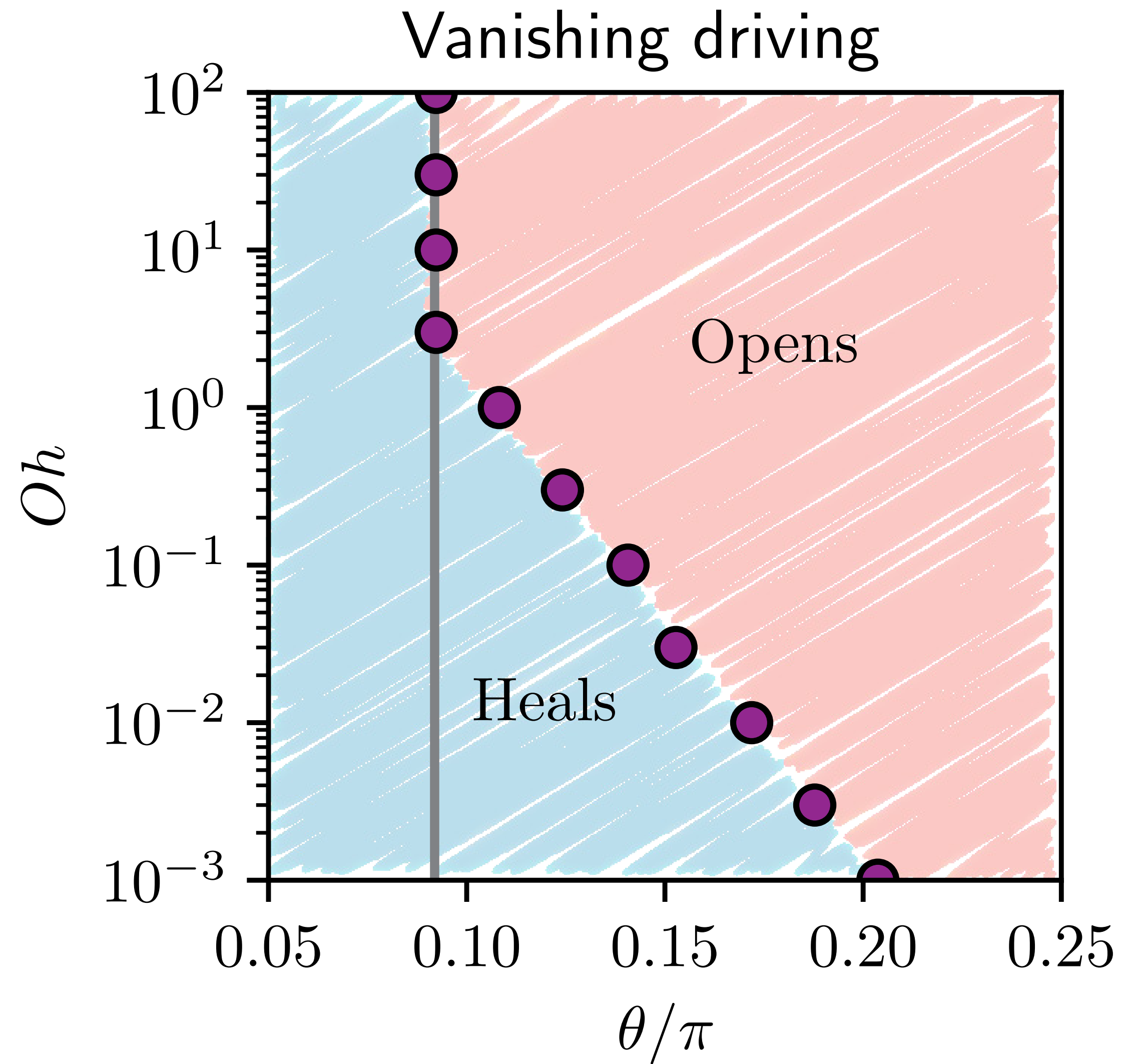
Taylor and Michael's criterion

$$\frac{R}{h_0} > 0.28$$

For vanishing inertia ($Oh \gg 1$)

$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

Summary



$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

Vanishing viscosity

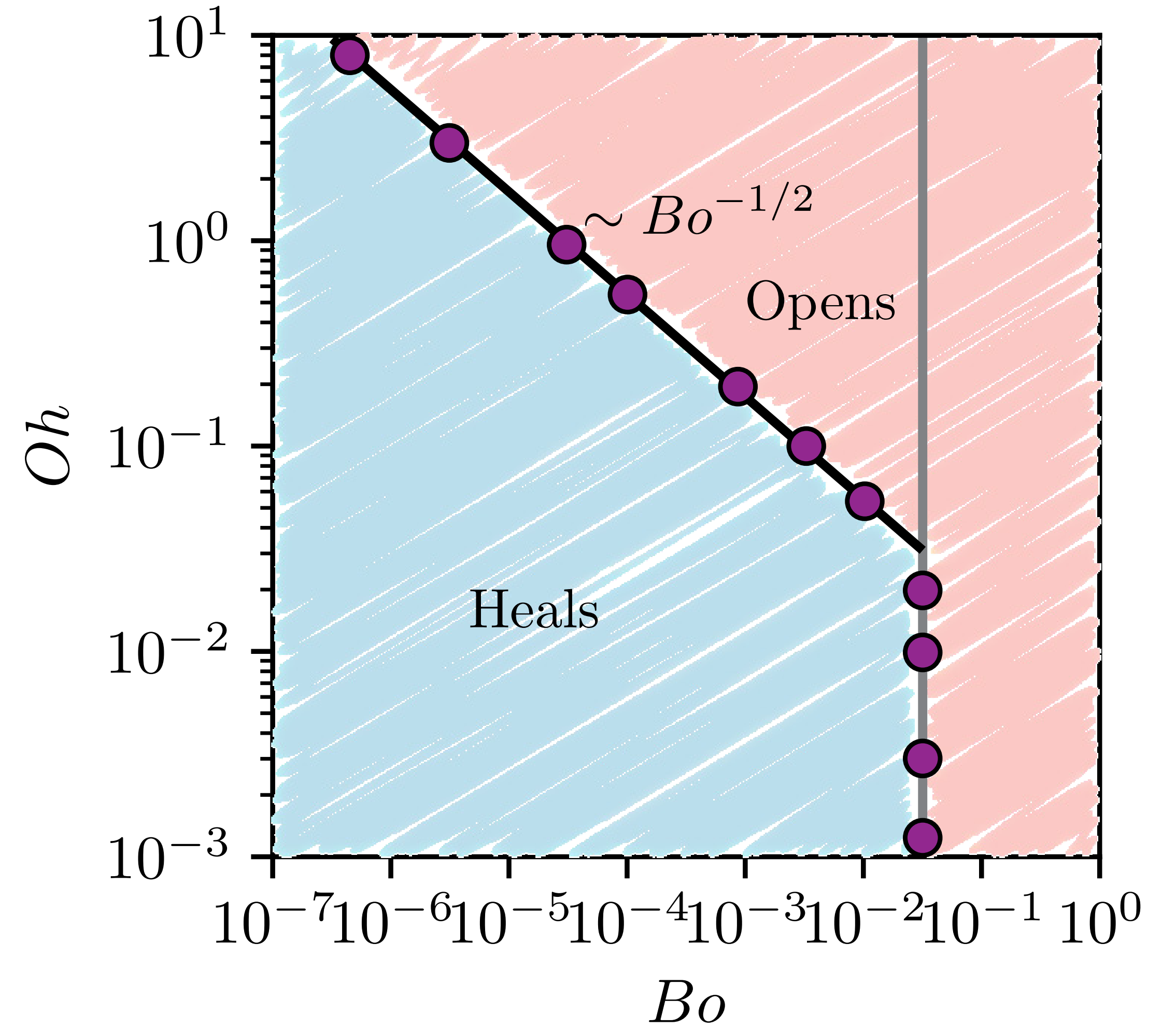
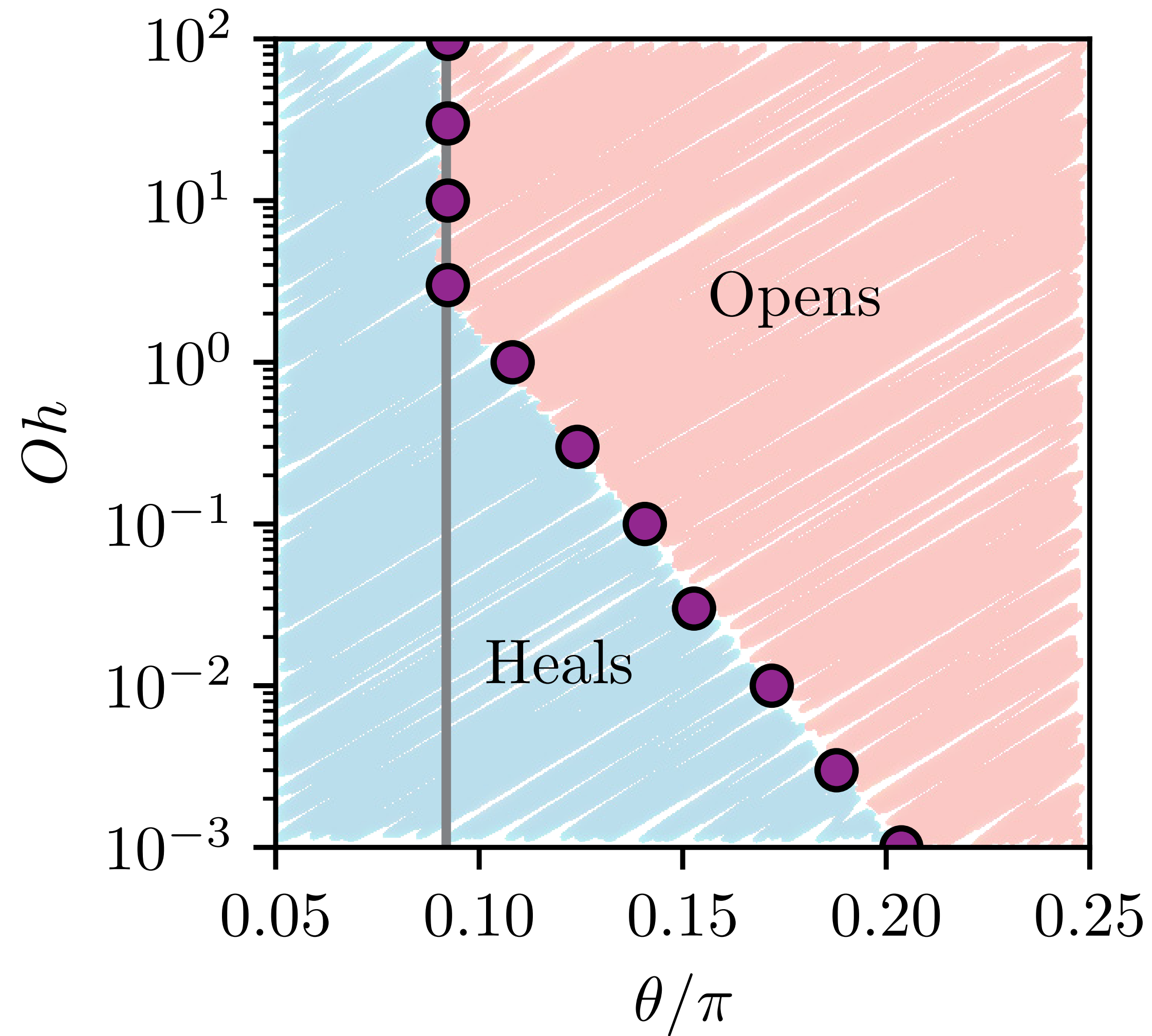
Double threshold behavior for breakup of liquid sheets

Detlef Lohse^{a,1}  and Emmanuel Villermaux^b 

1. Strong enough driving
2. Large enough distortions

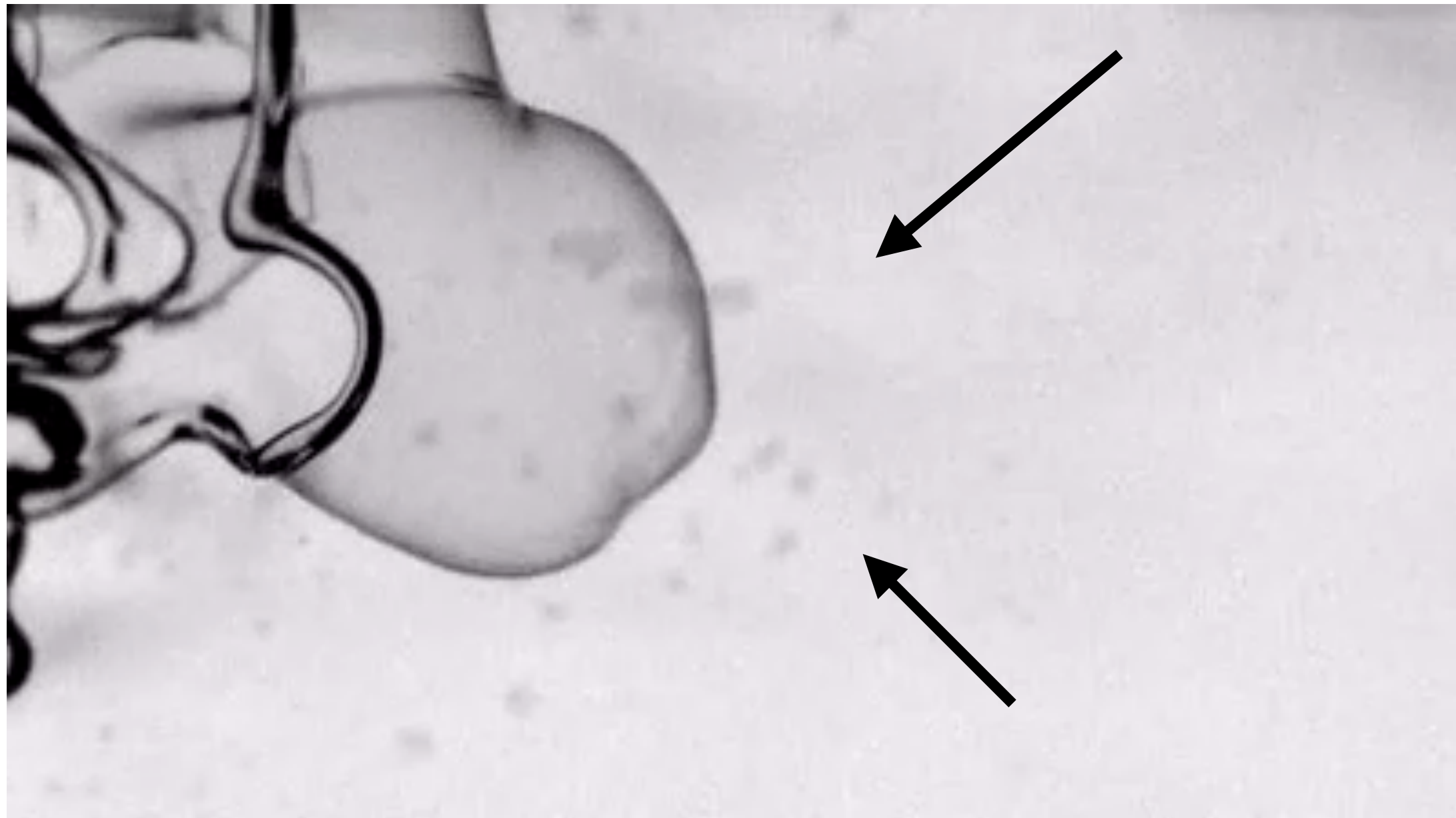
$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

Summary



$$Bo = \frac{\rho \omega^2 R_0^3}{\gamma} \quad Oh = \frac{\eta}{\sqrt{\rho R_0 \gamma}}$$

How do holes nucleate in sheets?

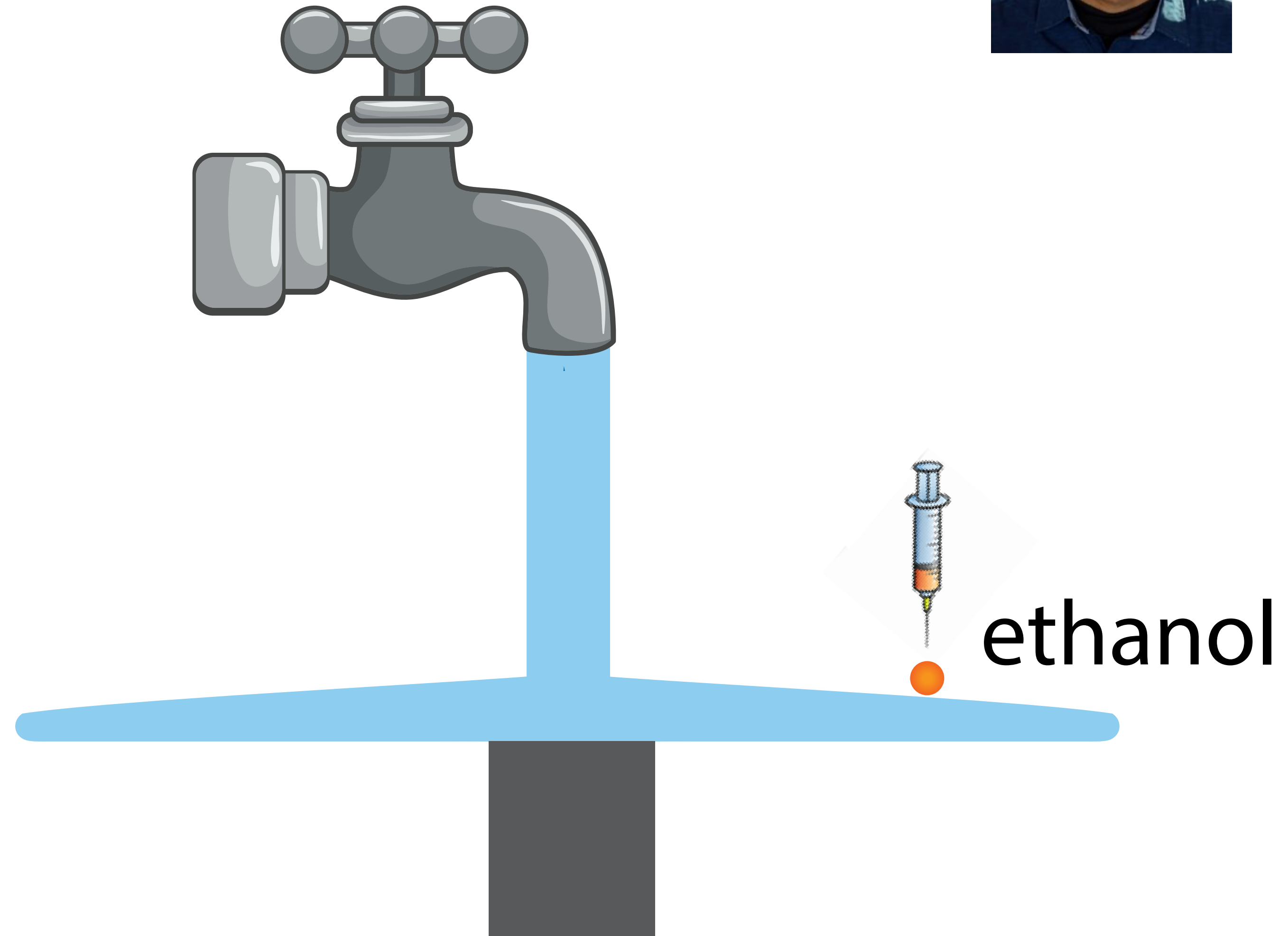


- Thermal fluctuations
- Van der Waal forces
- By hydrophobic particles or bubbles
- Chemical and temperature inhomogeneities

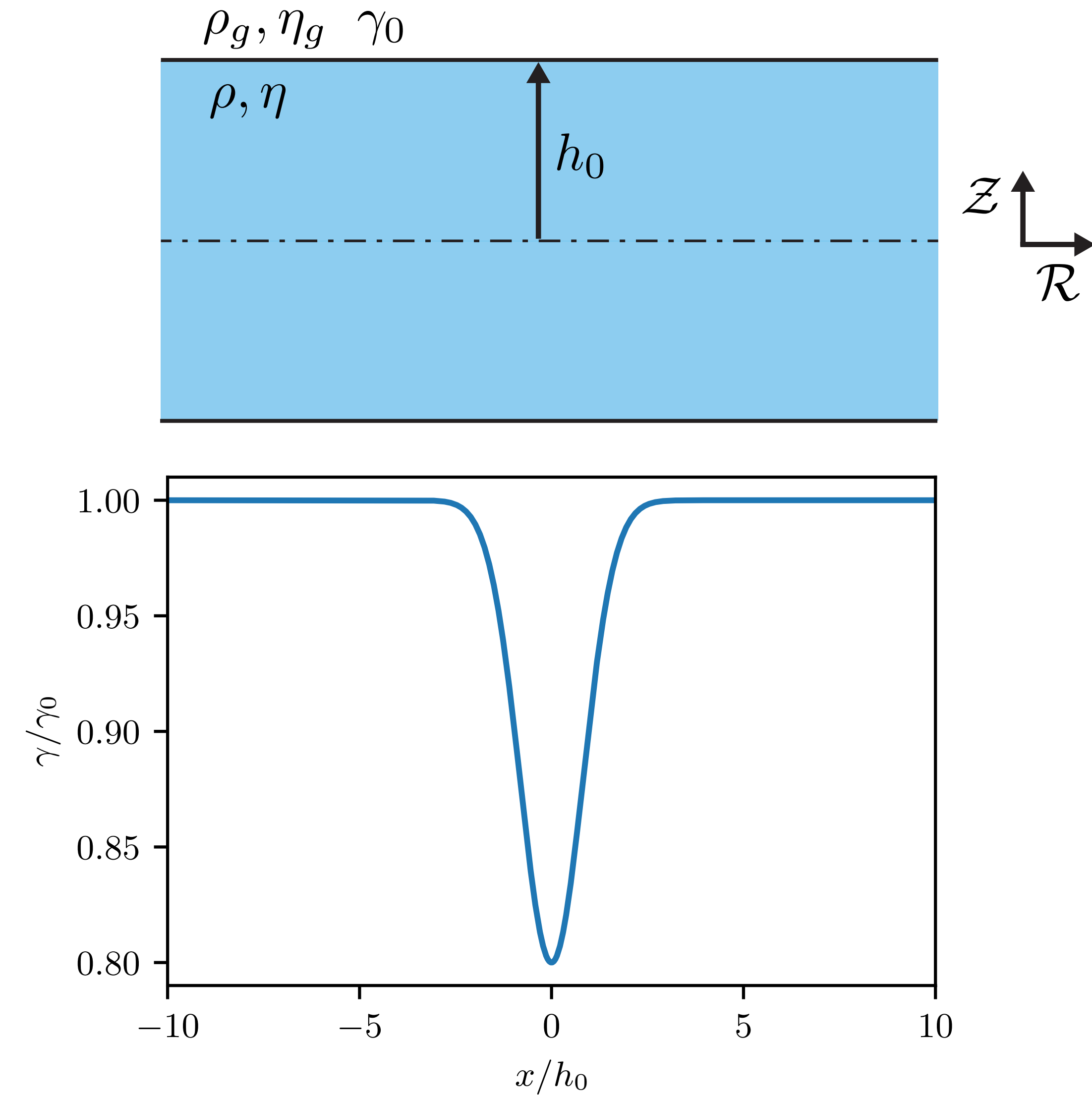
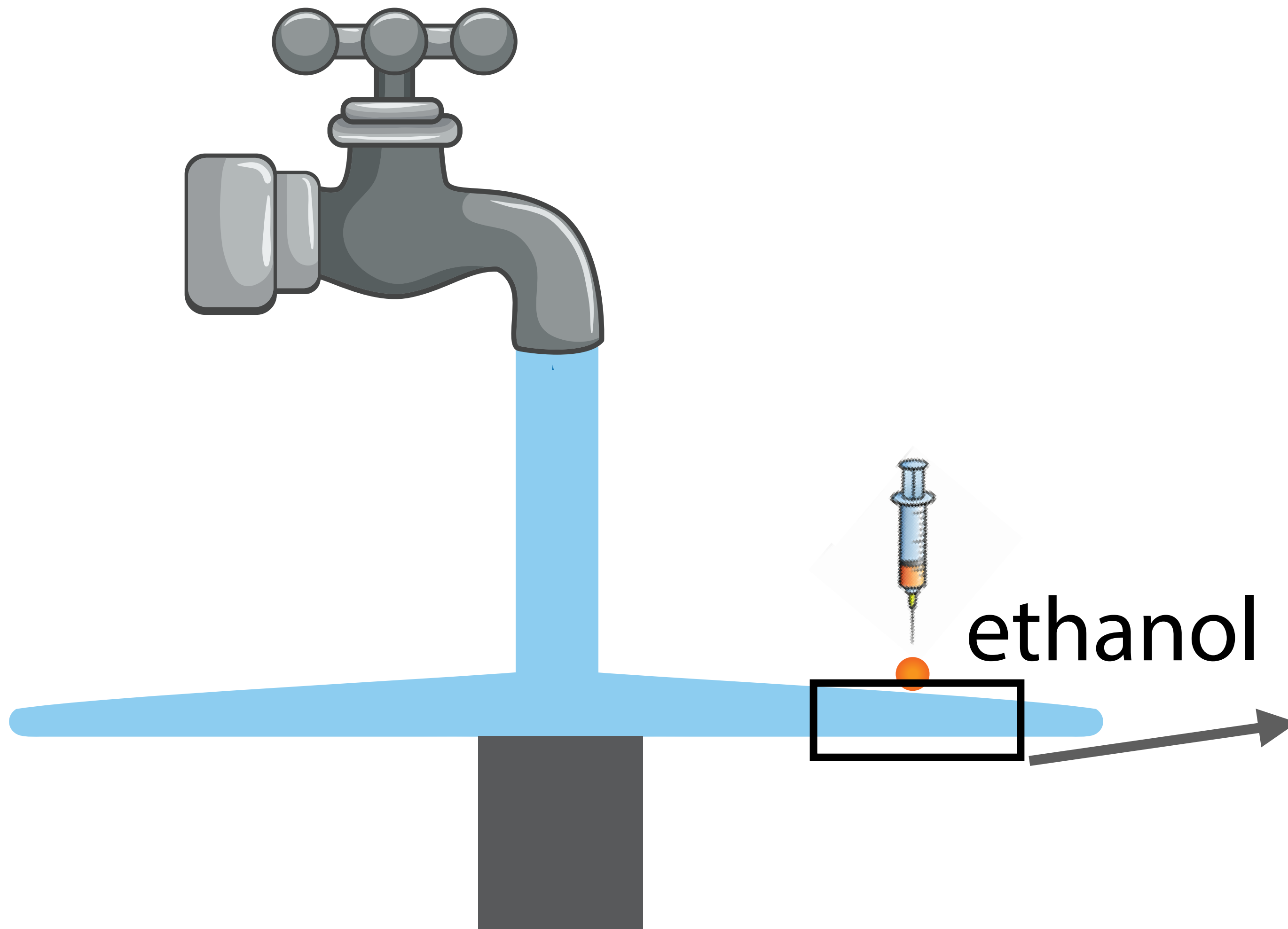
Open tap on a spoon



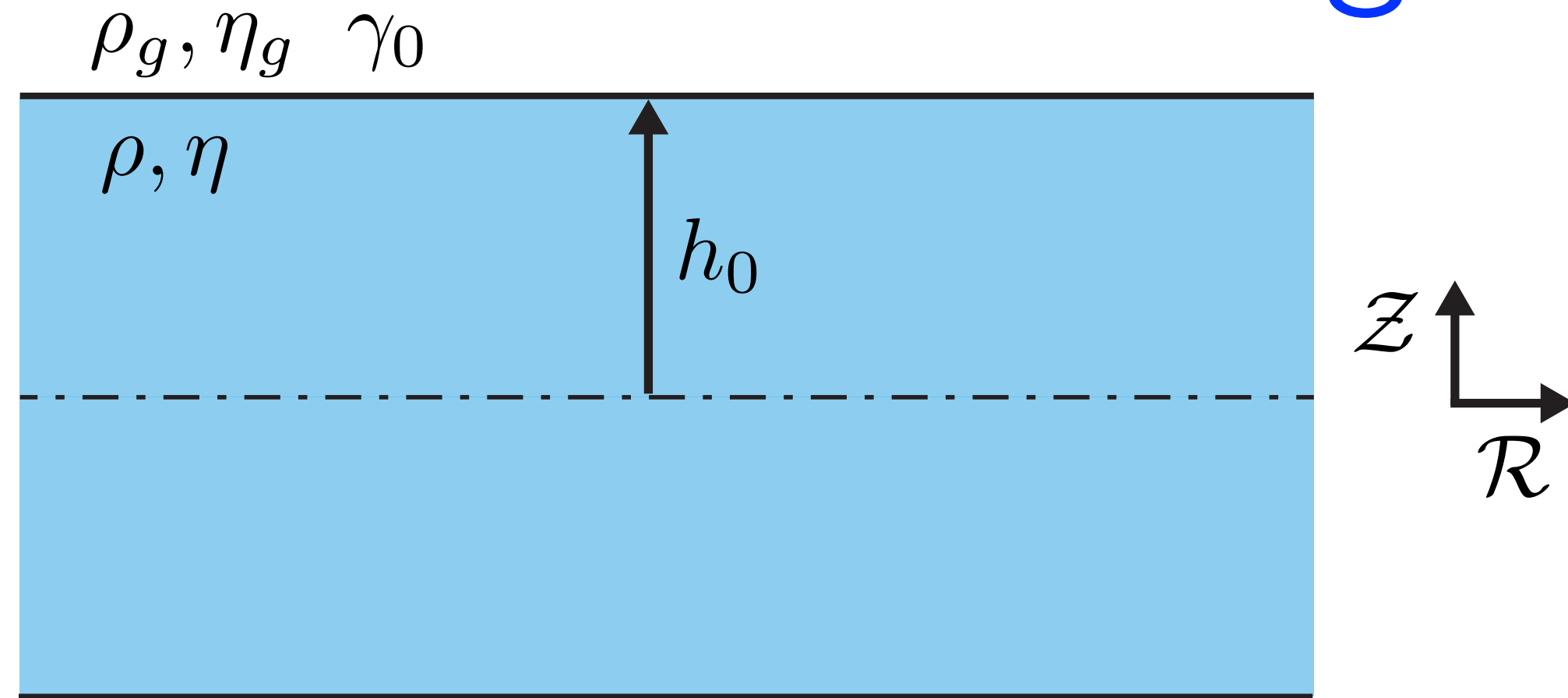
@r/WtWFotMJtRAtCaB



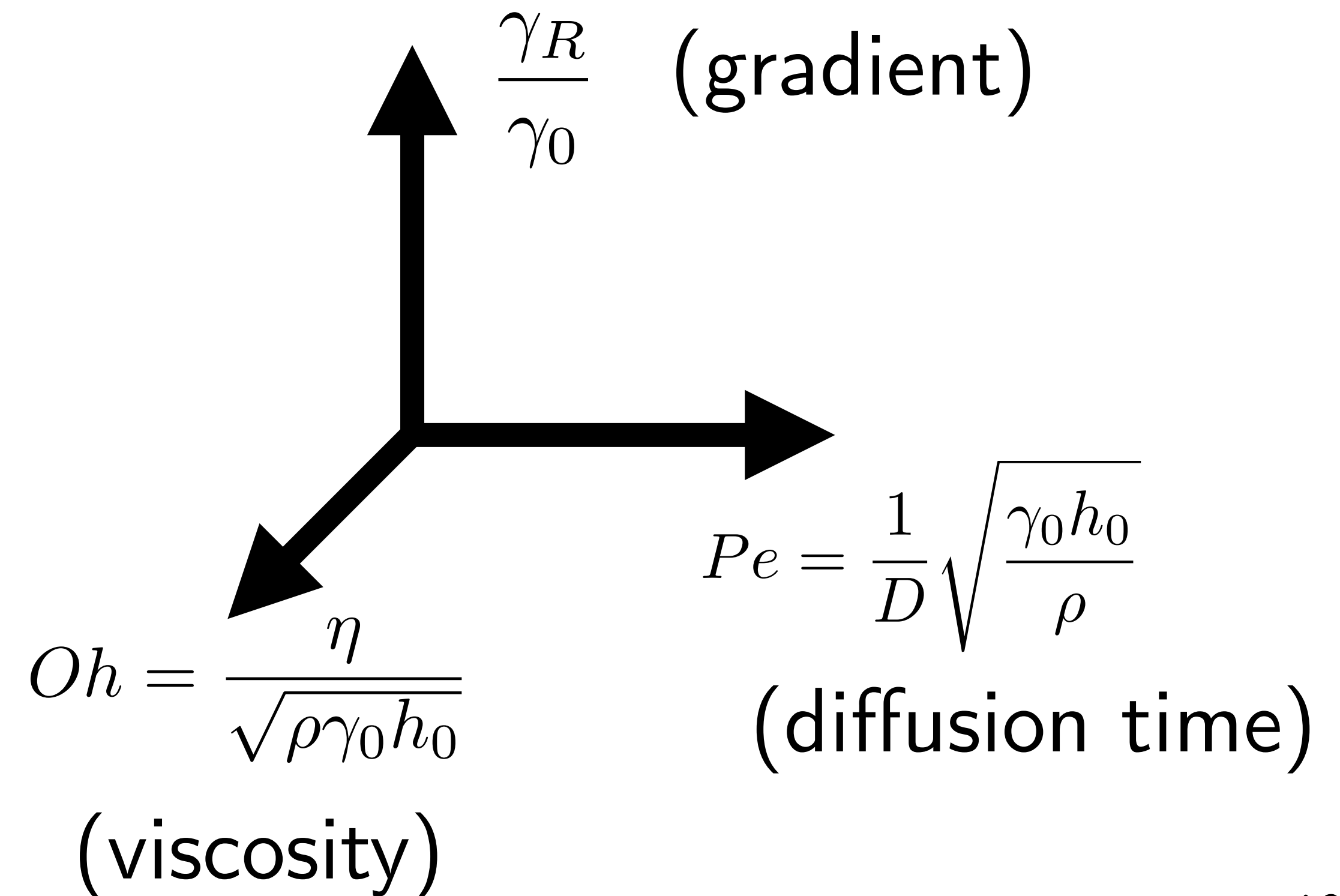
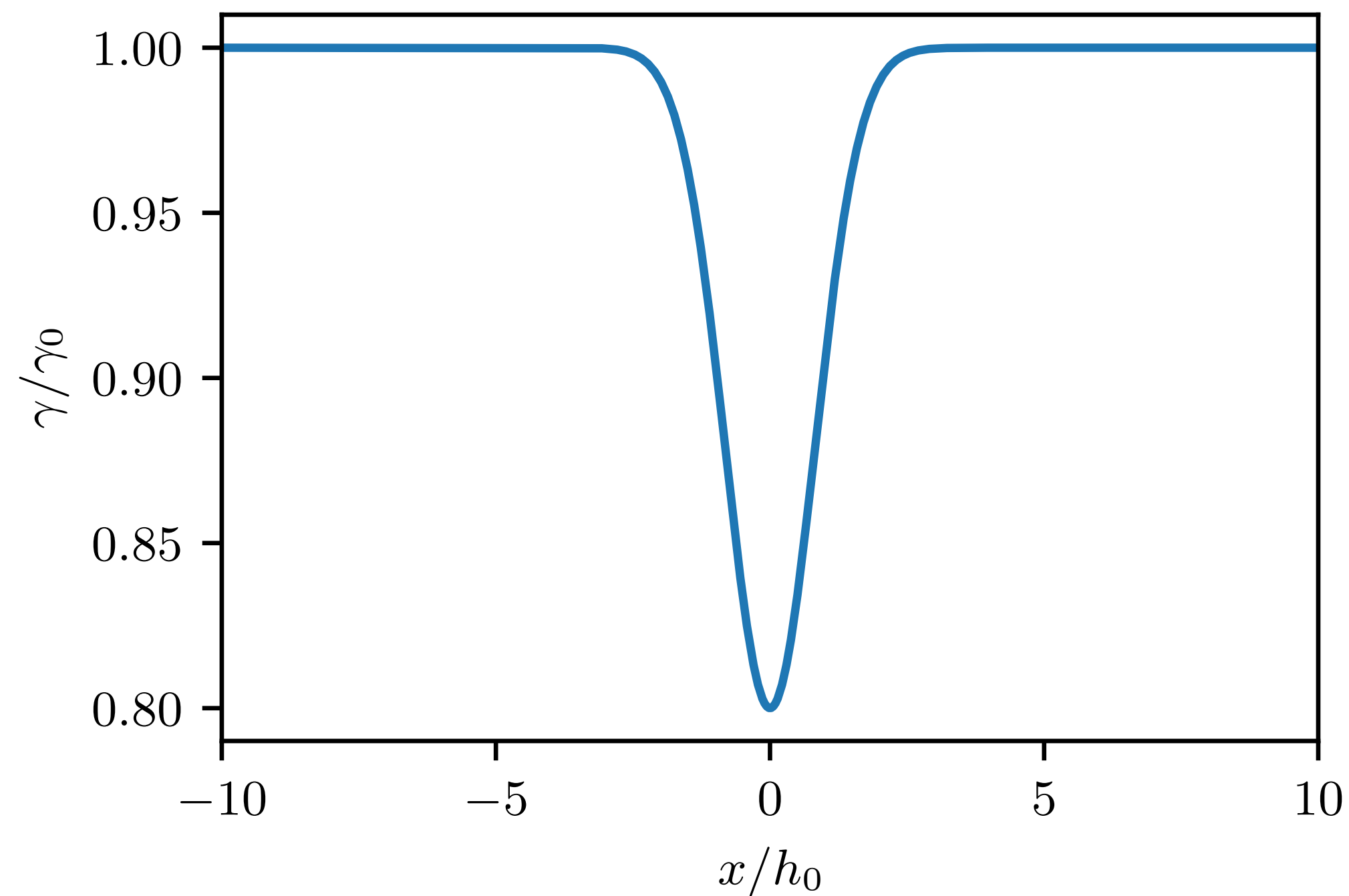
Open tap on a spoon



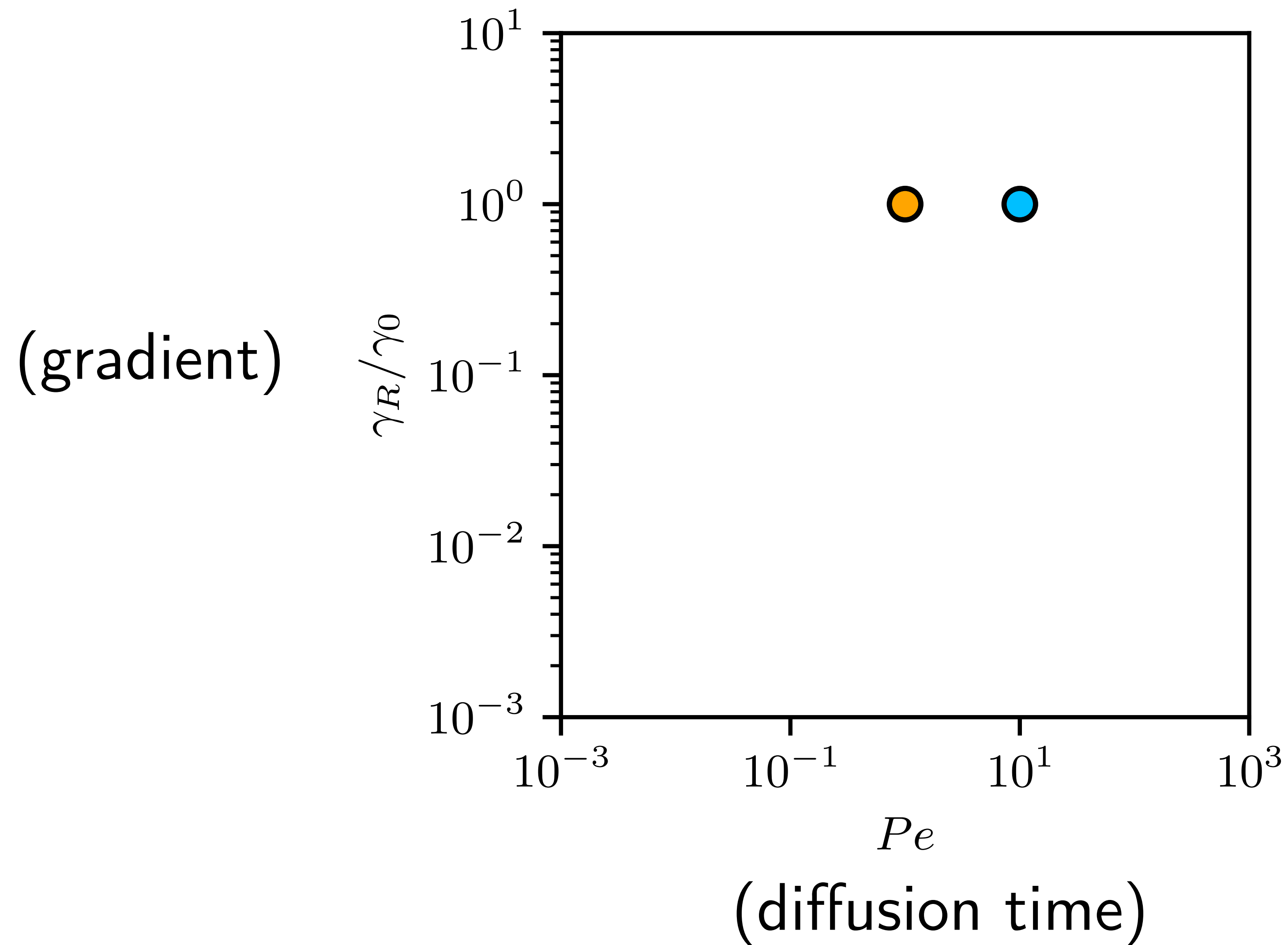
Marangoni driven sheet



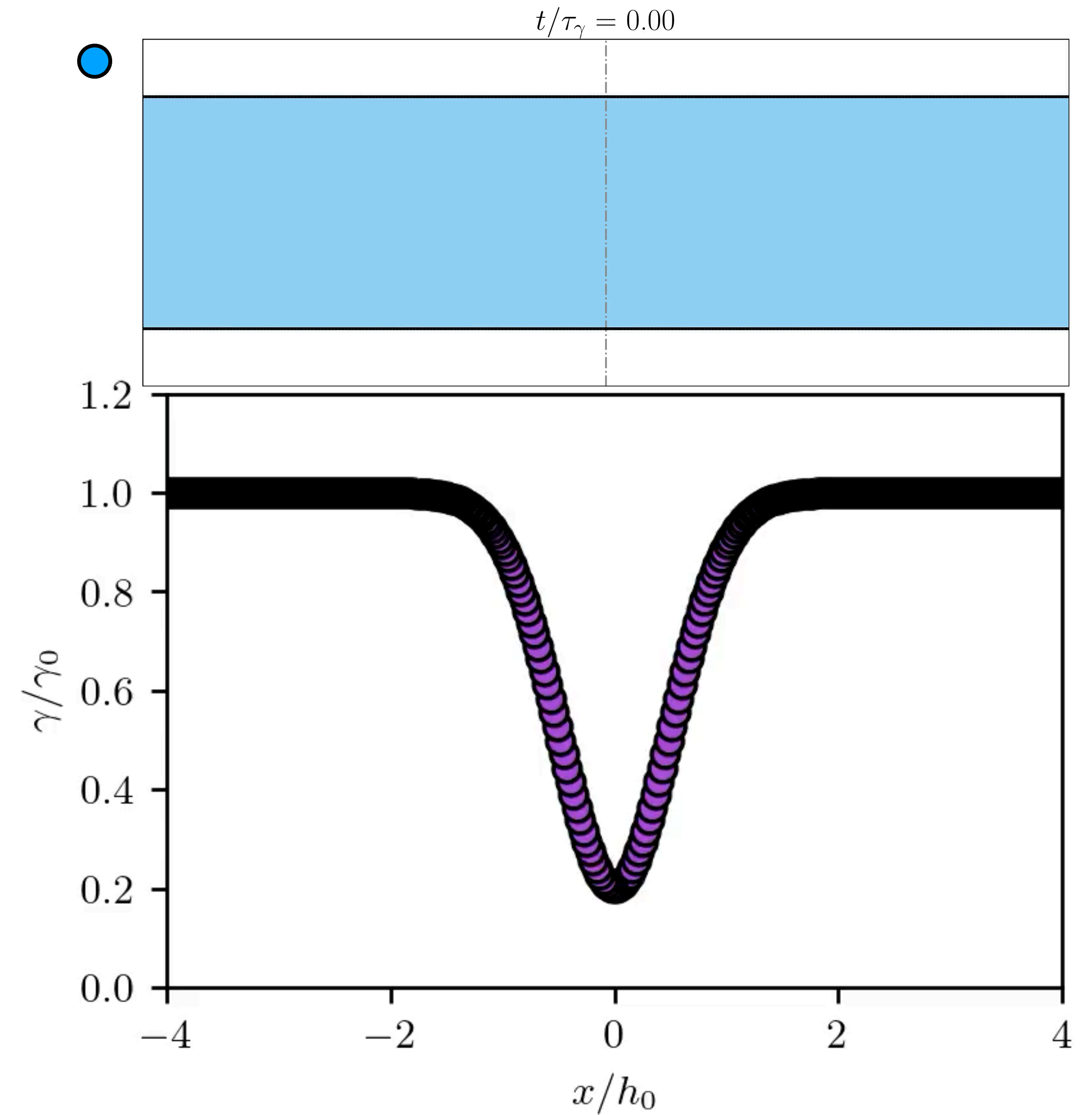
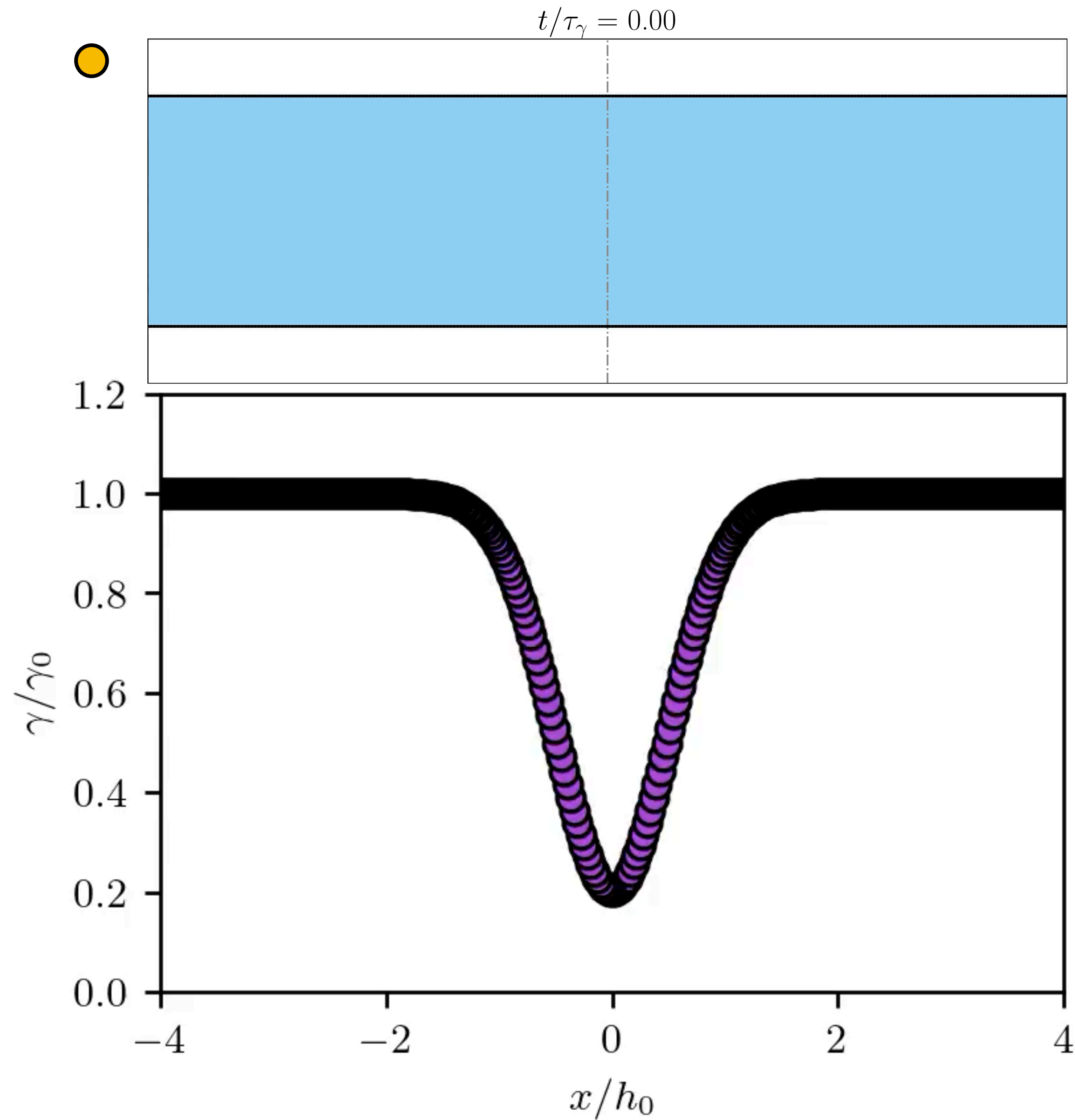
Finite D: $\gamma = \gamma_0 - \gamma_R c$



Parameter space



Effect of Péclet number



Thank you for your attention! Questions?

